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- Blazars in the Global Astrometric Instrument for Astrophysics (GAIA) era
- A case study: the optical counterparts of ICRF2 (and a future extension)
- The outcome from 2-D image decomposition of quasi-'stars'
- Ongoing plans & Summary

MOTIVATION

<u>Abstrometry</u> [the measurement of positions] is fundamental to, and the basis for, all other fields of Abstronomy. Indeed, even our "holidays" are based on astrometry...

- The forthcoming Gaia ESA mission will observe the solar system, the galaxy, and extragalactic objects, of which at least 500 000 are QSOs.
- Objects with core radio morphology, absent proper motions, apparent pointlike nature ensure a high degree of accuracy and stability of their coordinates.
- These kind of objects can help on the alignment between the submilliarcsecond radio frame (ICRF) with the existing and emerging optical reference frames of similar accuracy, constructed by astrometric satellites (GCRF).
- Our aim is to study one of the main difficulties in achieving this goal: the extended structures of quasar hosts.

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We performed a case study by finding optical counterparts to the ICRF 2 radio defining sources (and future extension -Bourda et al, 2011) in the SDSS DR 9.

- ICRF2 uses a set of <u>295</u> extragalactic sources distributed over the entire sky and selected on the basis of positional stability and the lack of extensive intrinsic source structure. The precision of the source coordinates is better than <u>1 mas</u>.
- A sample of <u>105</u> optically-bright extragalactic radio sources as compiled (Bourda et al,2011) by cross-correlating optical and radio catalogs and in order to upgrade the defining sources in the current reference frame. The precision is around <<u>200µas</u>.

From our sample of 400 sources, only 198 have optical counterparts in the Sloan Digital Sky Survey Data Release 9.



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July 2013

RESULTS



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 $R_e \ [kpc]$

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RESULTS

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<u>Gata Instrument and Bastc Stmulator ts a ptxel-level stmulator</u> intended to simulate how Gata instruments will observe the sky, using realistic simulations of the astronomical sources and of the

înstrumental properties.

Object apparent size (300mas)

We aim to analyze the ICRF2+ objects by asserting on their detectability through GIBIS and, if possible, inferring their morphological parameters as seen by GAIA simulated data.

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The combination of Gata optical astrometry and accurate radio astronomy will probe directly the geometry and physics of the targets as never done before, which holds promises for new discovertes...

The detailed analysis of the images can permits to pinpoint an extended component (host galaxy), which may be a source of error for centroid determination, if detected by Gaia.

Answer (for this sample): 9%. Guess (for now): +17%.

Thank you for your attention and have a nice lunch 🙂

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SUMMARY