

# Towards the discovery of the *first* radio galaxies





# Where are the first radio galaxies?

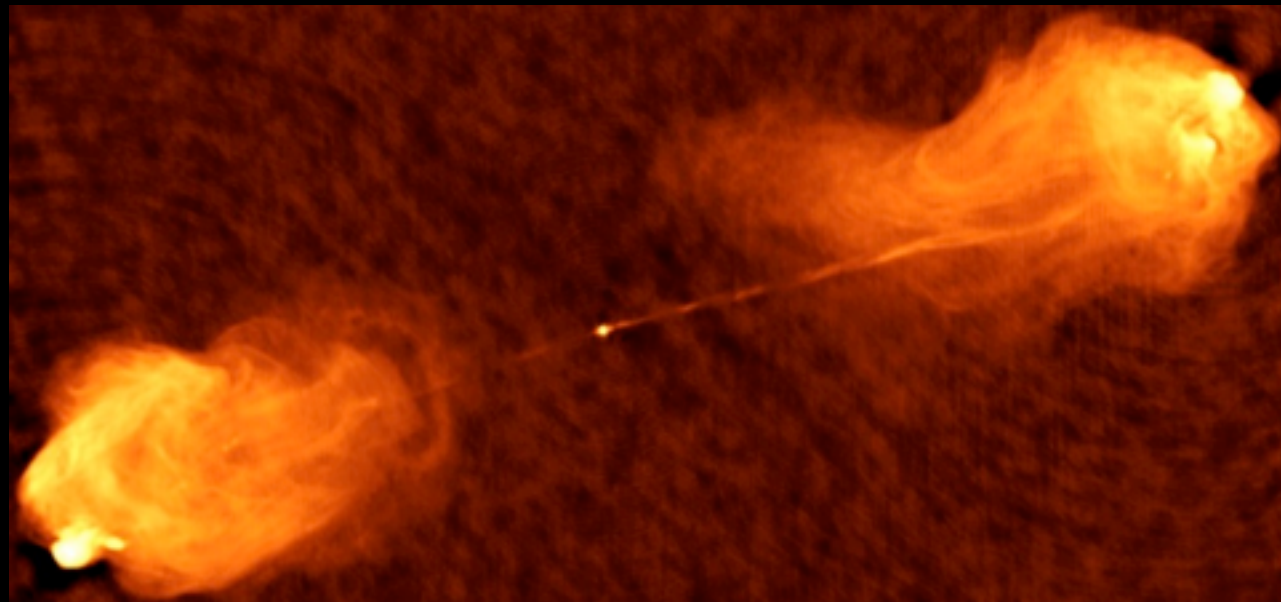
# Where are the first radio galaxies?

- The earliest AGN activity in the Universe.
- Powerful tracer of AGN physics at very high redshifts.
- If (when!) found will allow for HI absorption studies in the EoR.
- Possibly already *well* within current day capabilities.

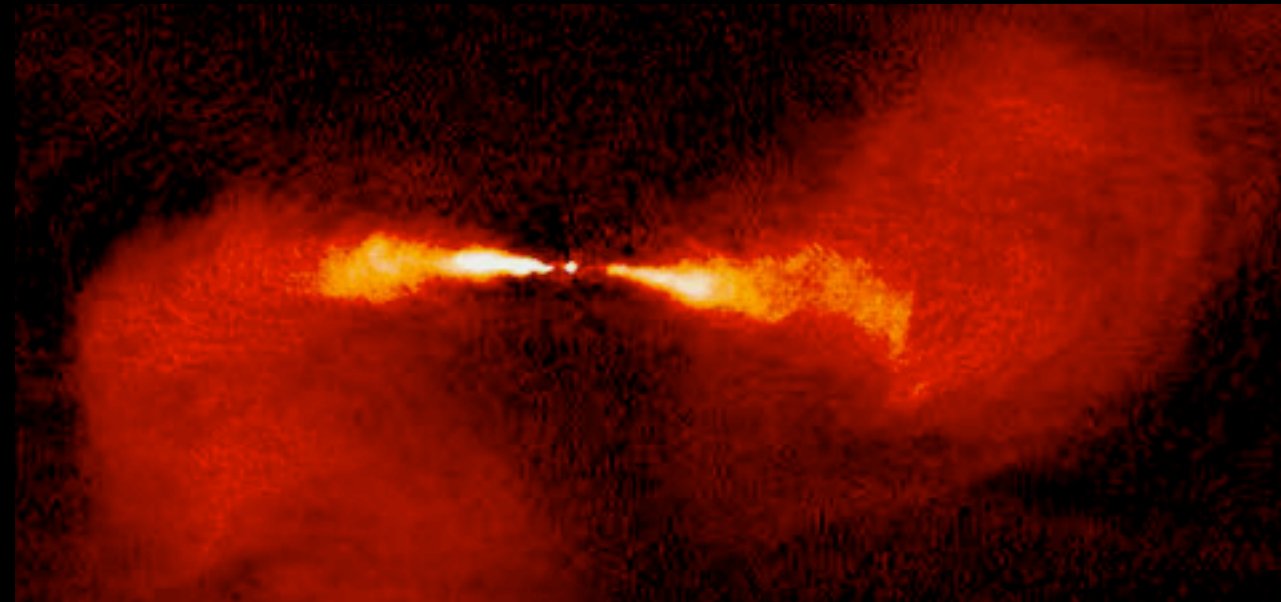


# Radio Galaxies = Active Galactic Nuclei

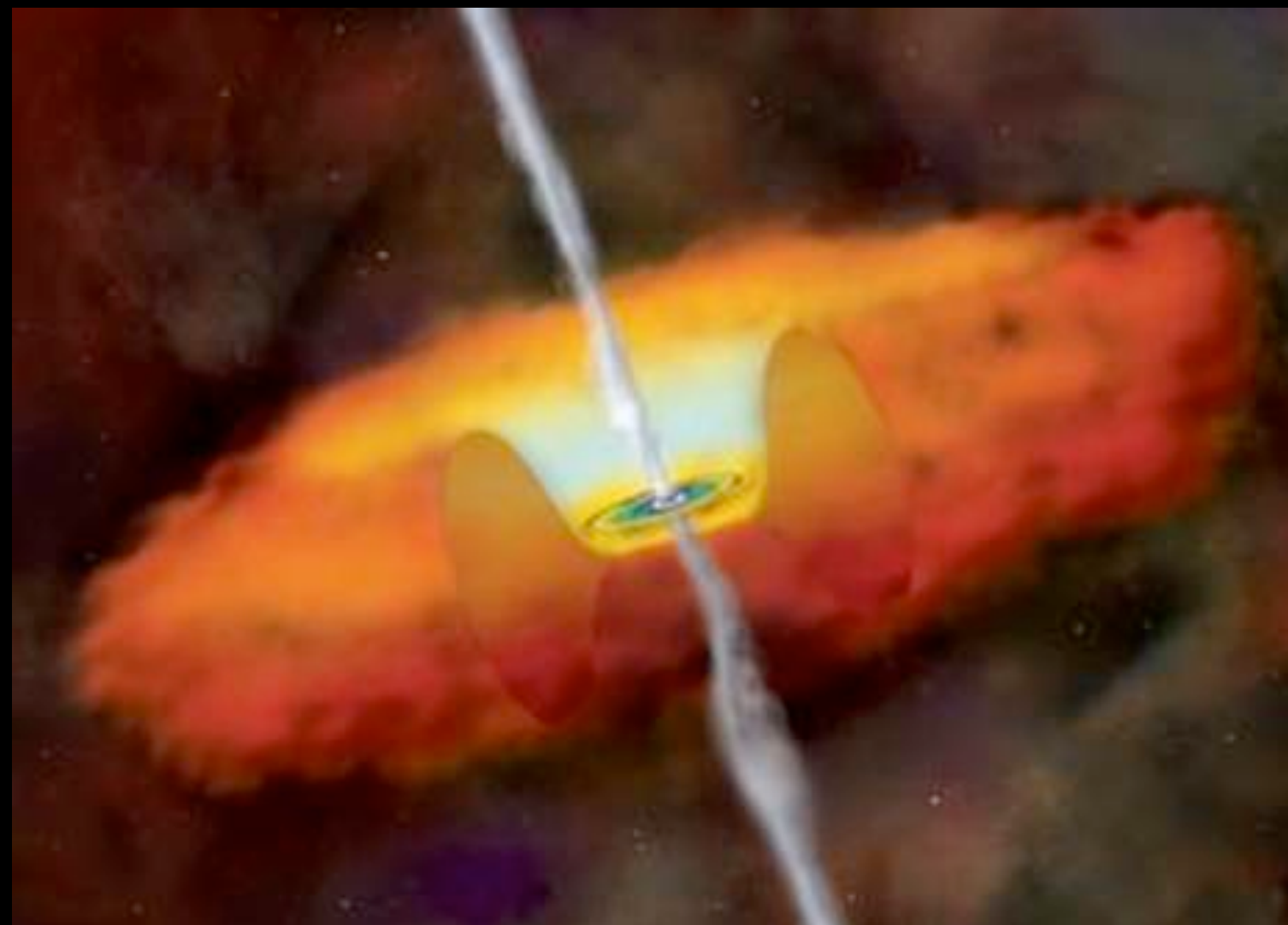
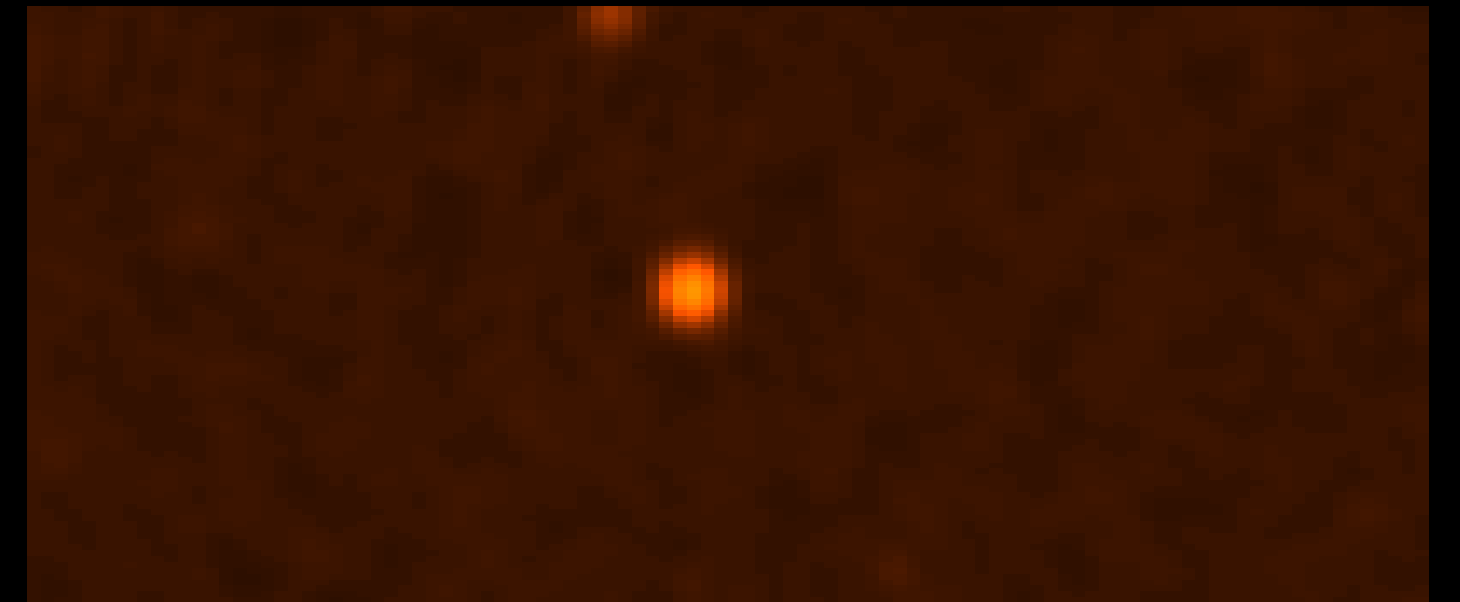
More powerful...



...less powerful...



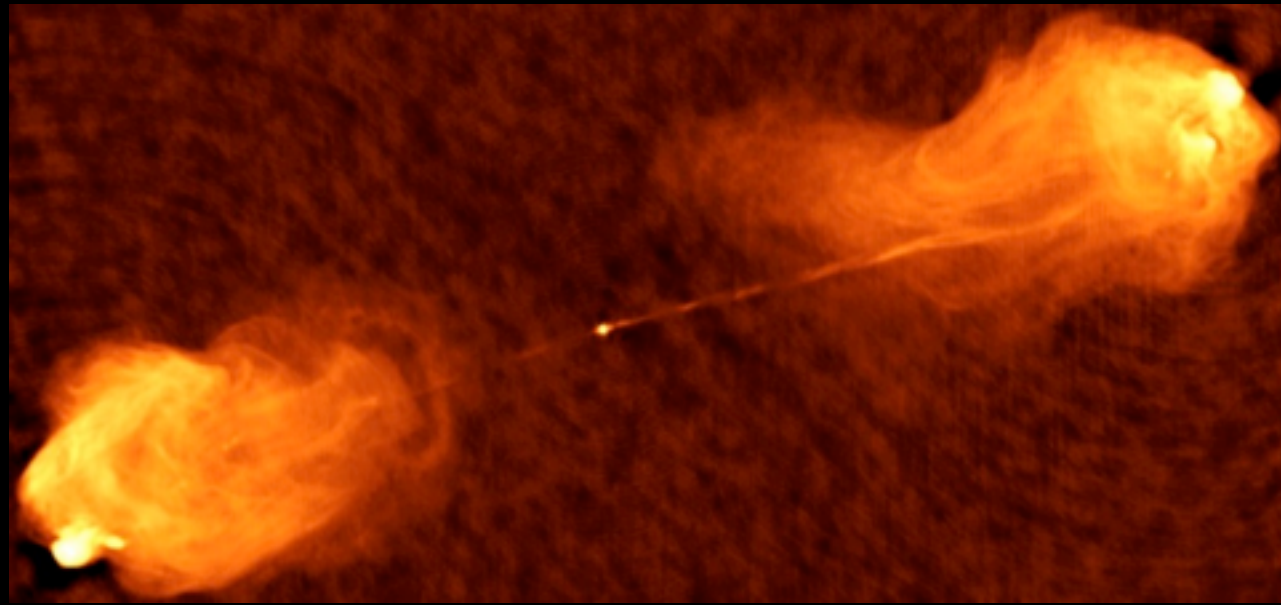
...or “radio-quiet”.





# Radio Galaxies = Active Galactic Nuclei

More powerful...



- 1) Among the most luminous galaxies at any redshift
- 2) Associated with the most massive systems
- 3) Progenitors of brightest cluster ellipticals
- 4) Track proto-cluster environments
- 5) Dusty and Starbursting
- 6) May show large gas reservoirs ( $\text{Ly-}\alpha$  halos)

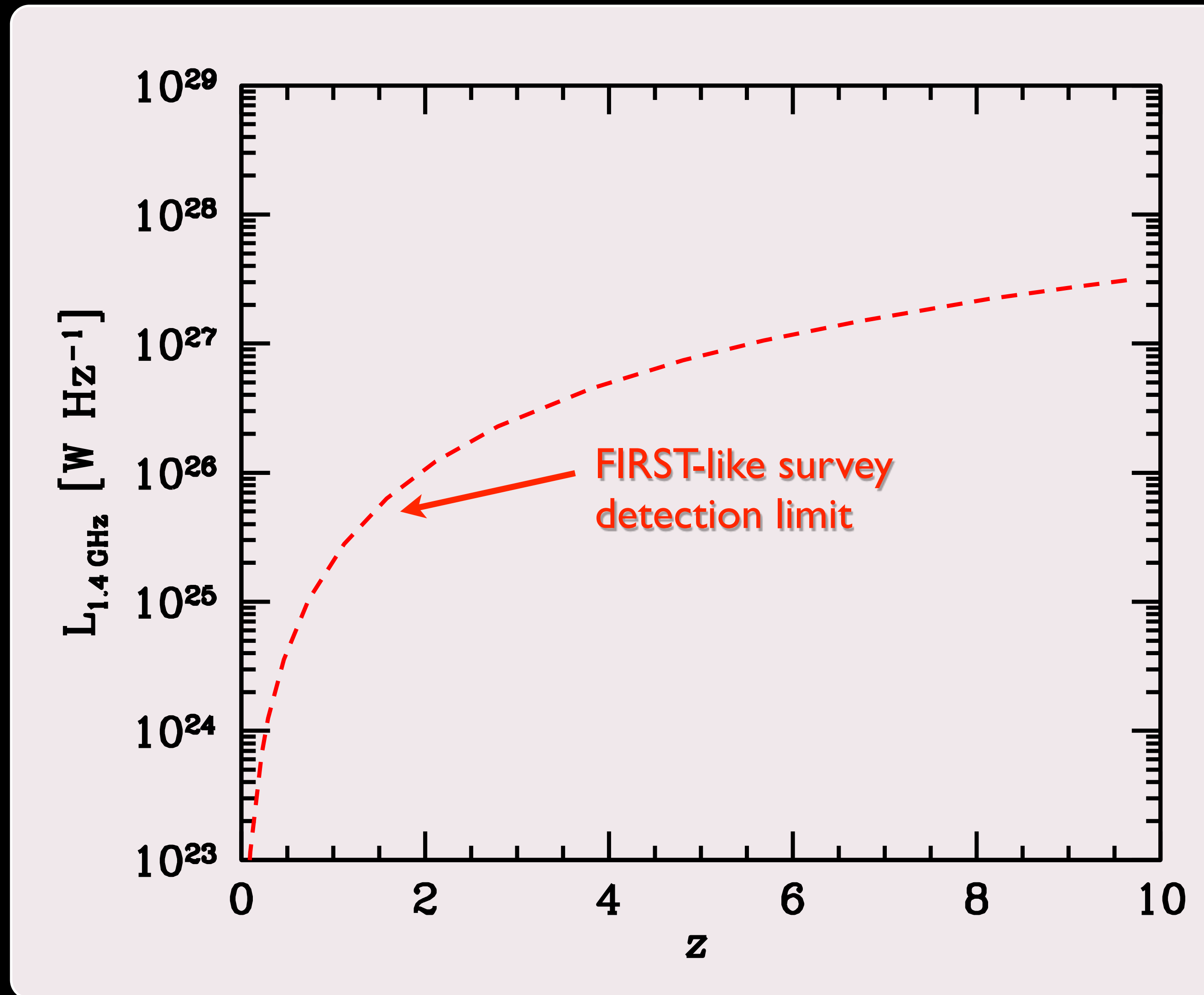
Tracers of galaxy buildup AND structure formation ( $z > 2$ )



# The (radio) road to the highest redshifts

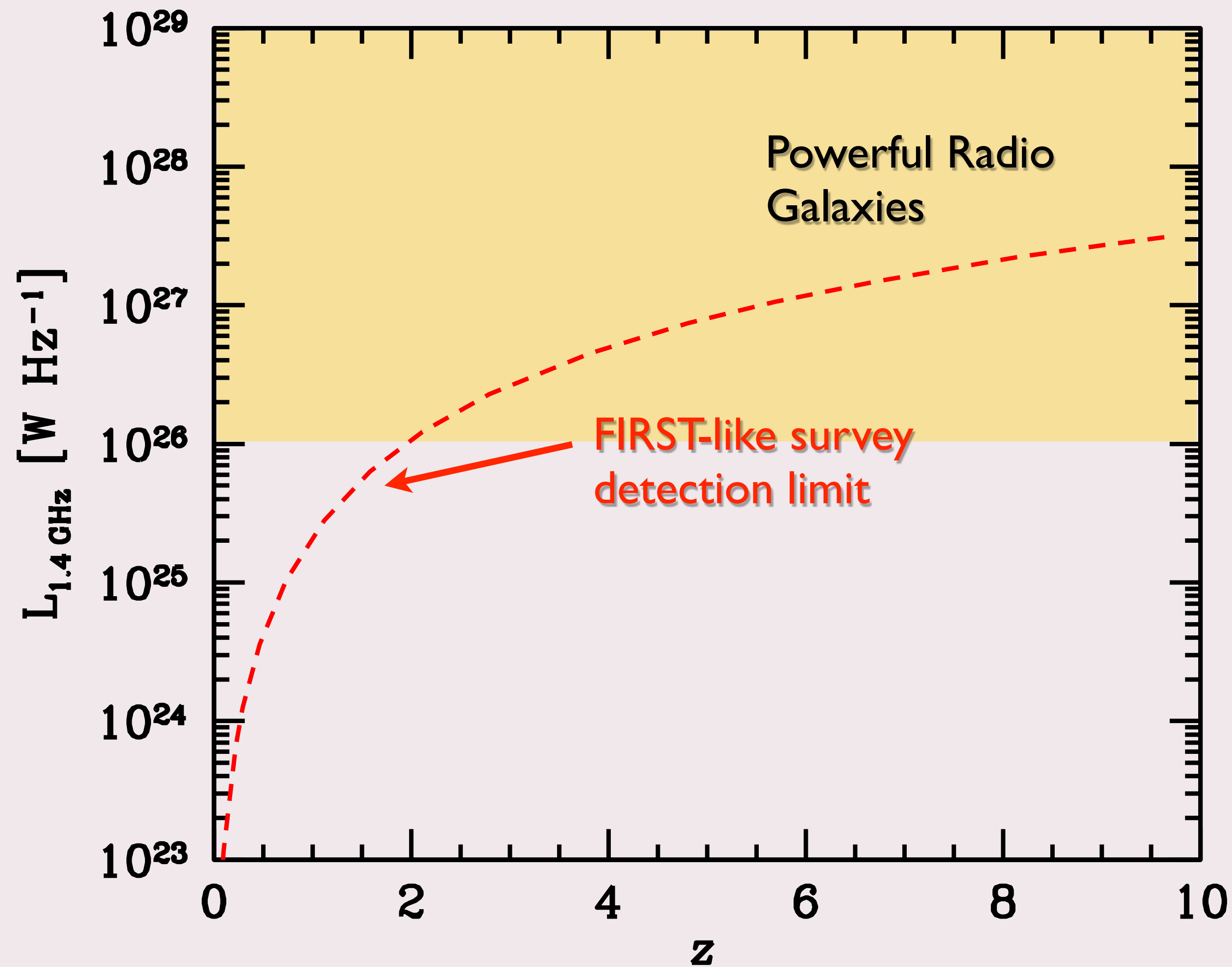


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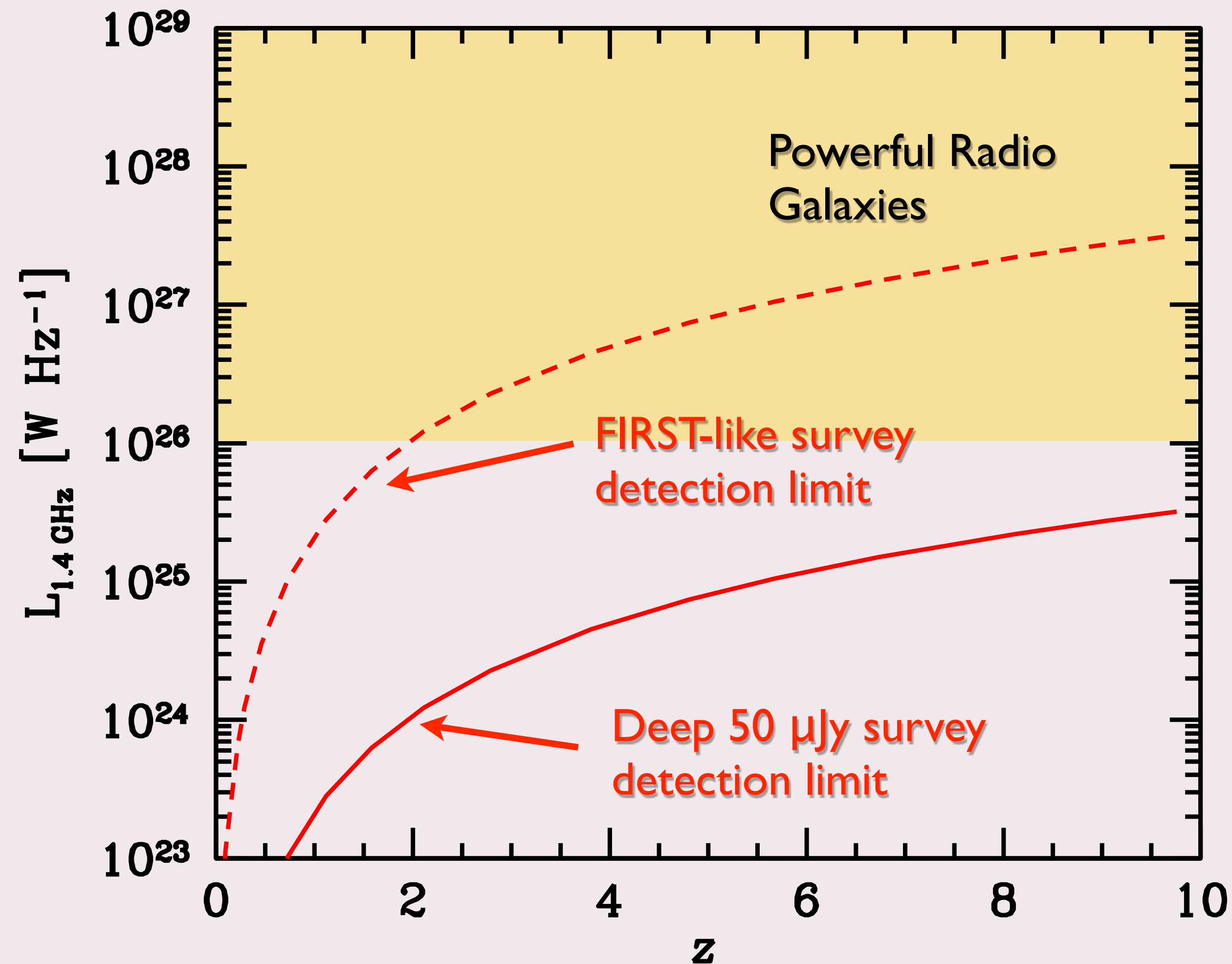


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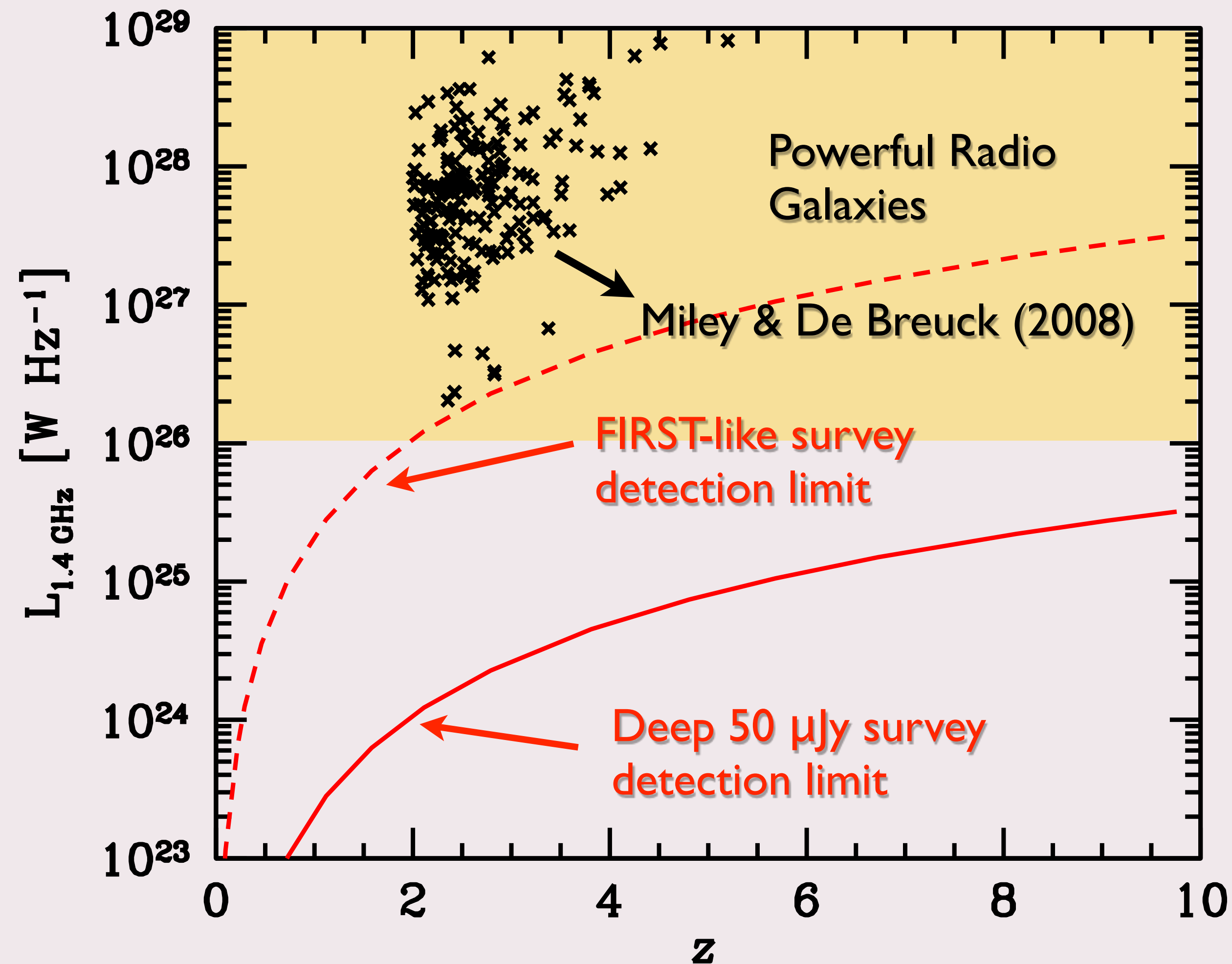


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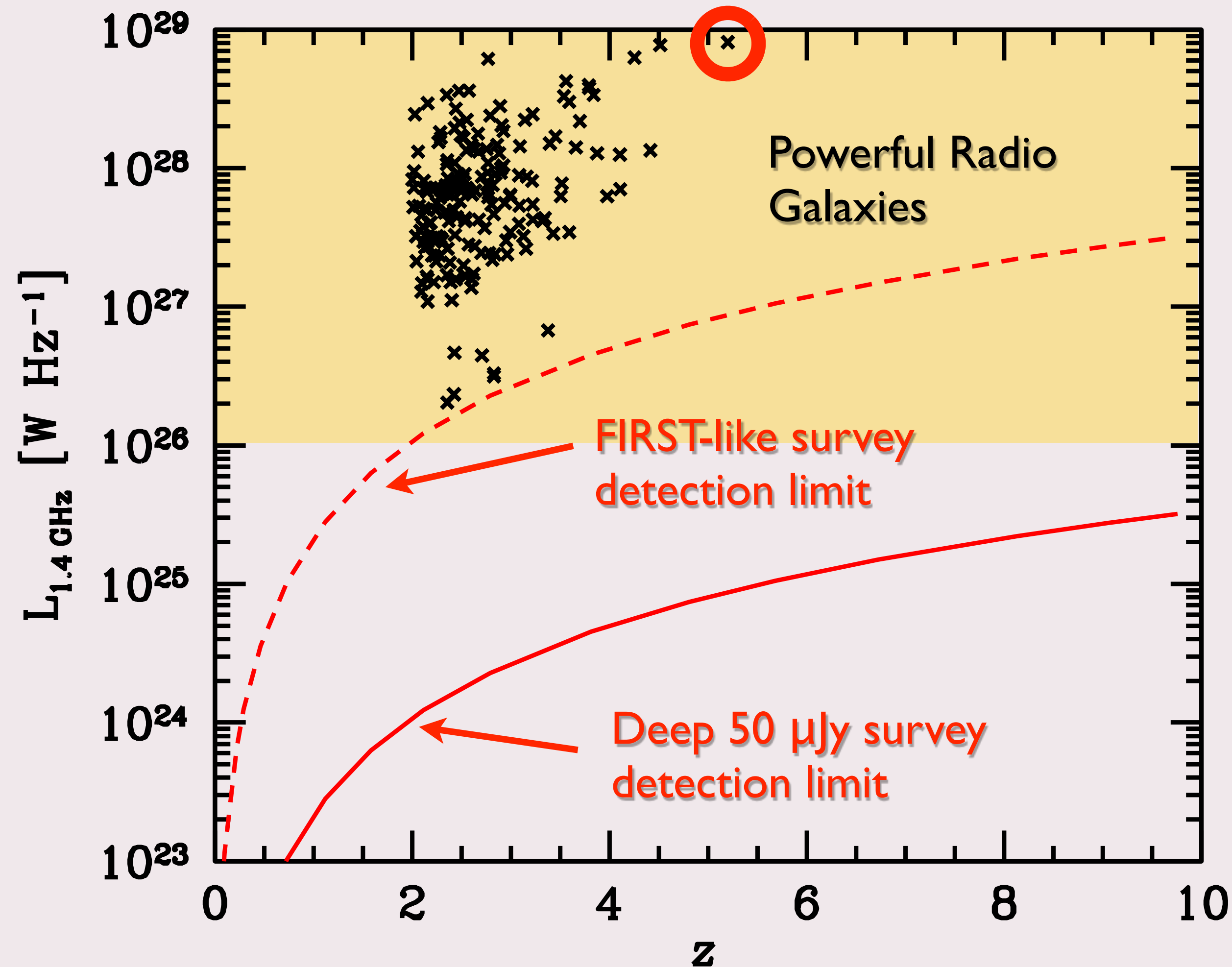


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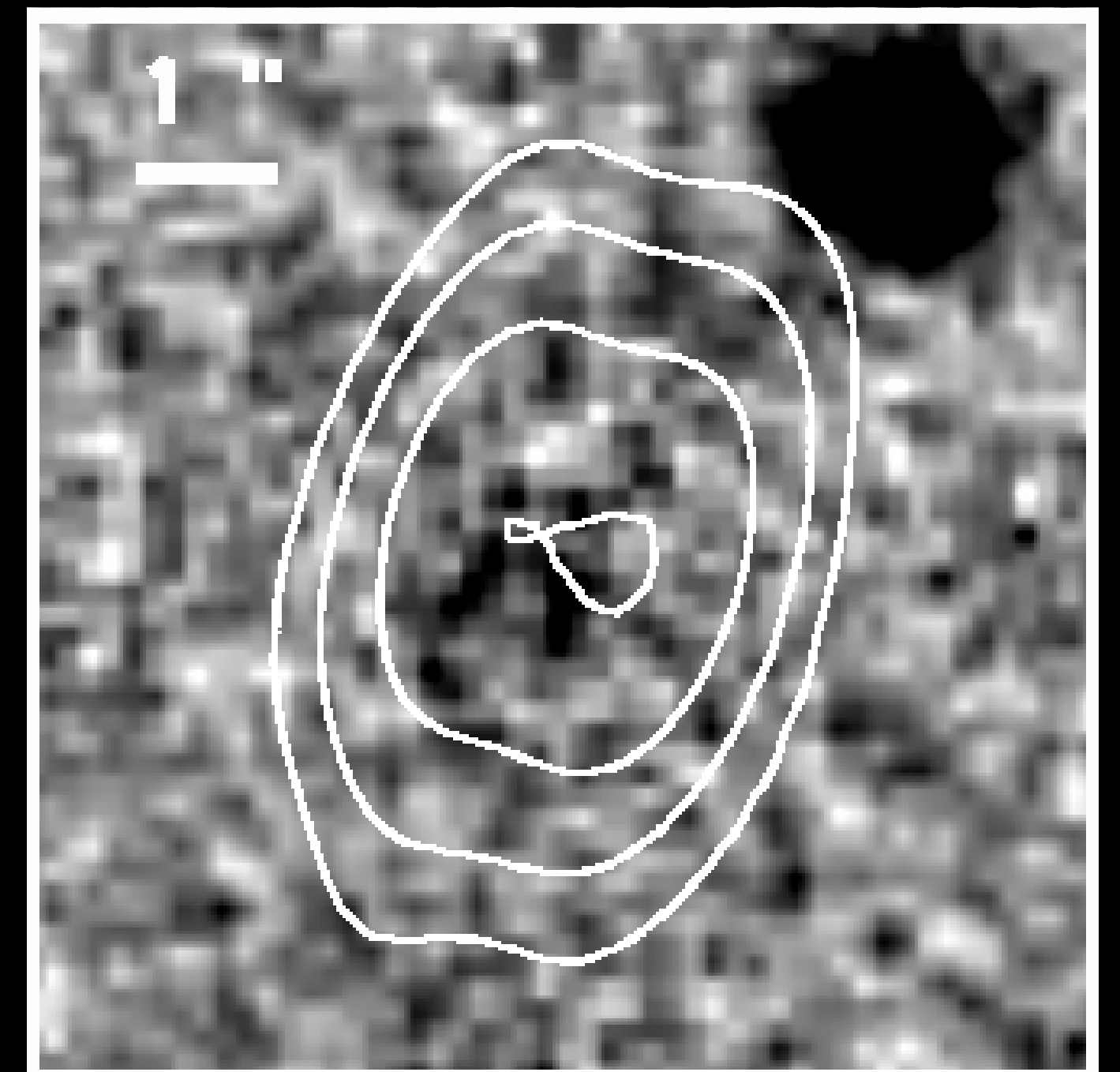




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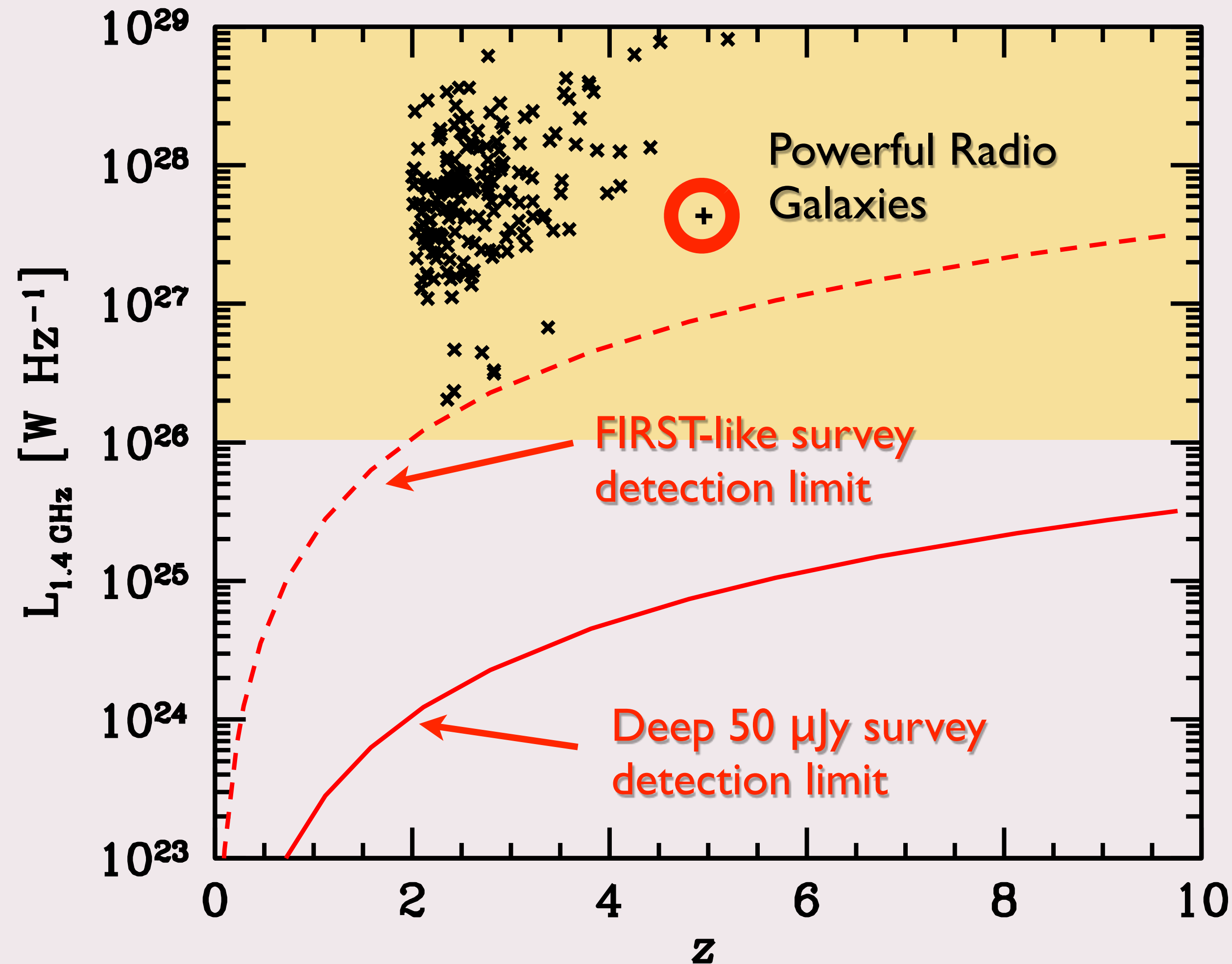
TN J0924-2201 @  $z=5.2$   
(van Breugel+, 99)



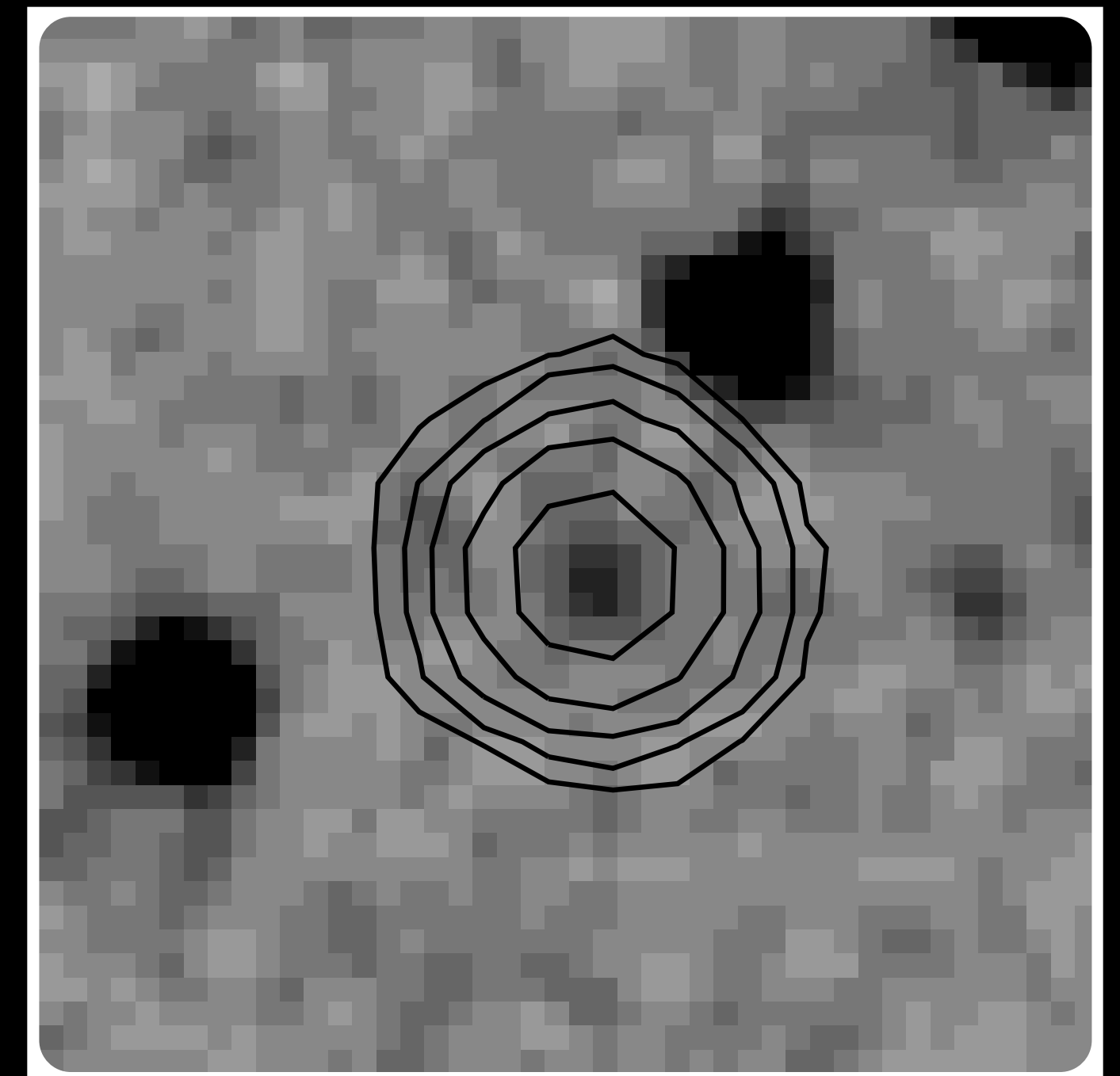
$S_{1.4\text{GHz}} = 73 \text{ mJy}$



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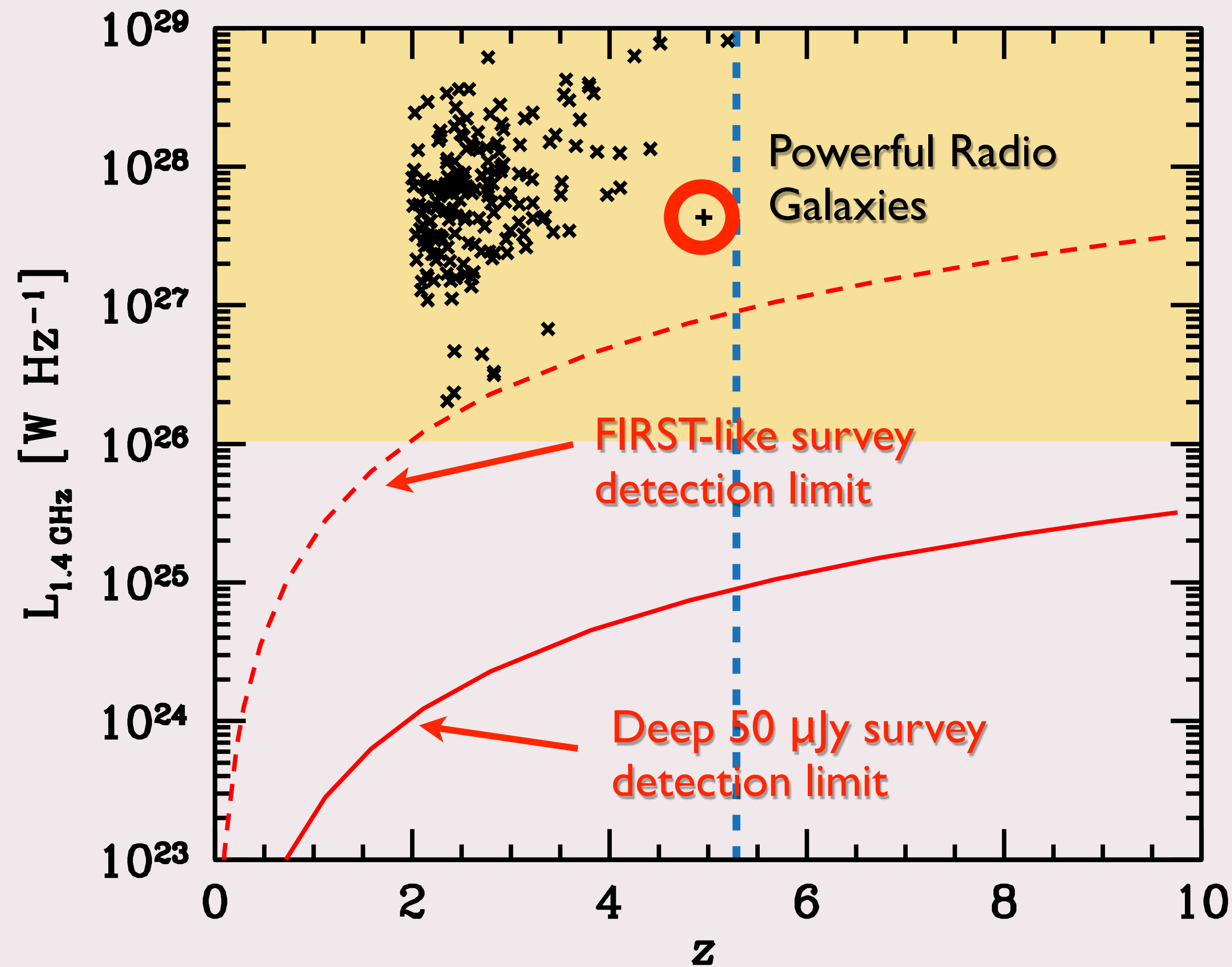
J163912.11 + 405236.5 @  $z=4.9$   
(Jarvis+, 09)



$S_{1.4\text{GHz}} = 22 \text{ mJy}$

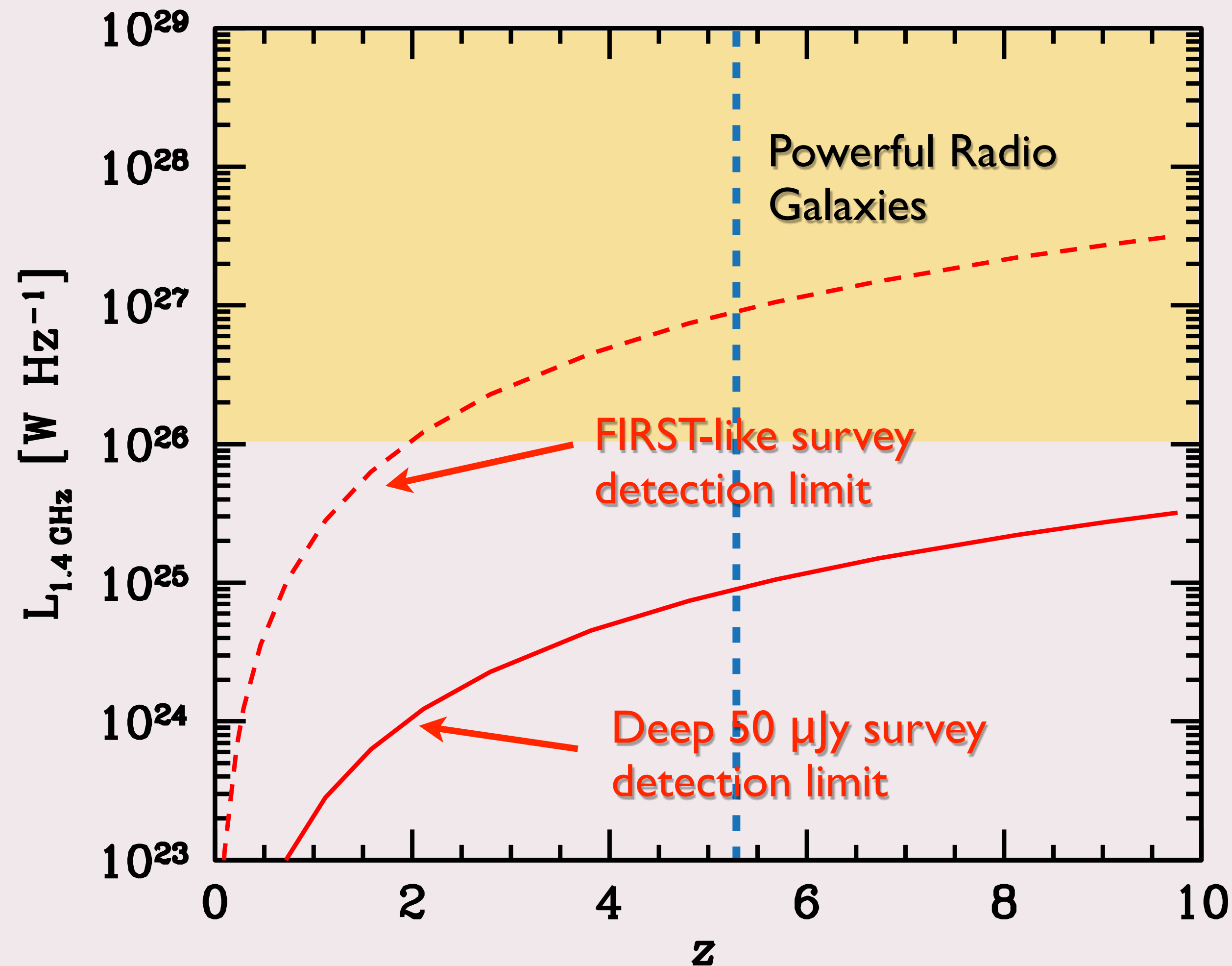


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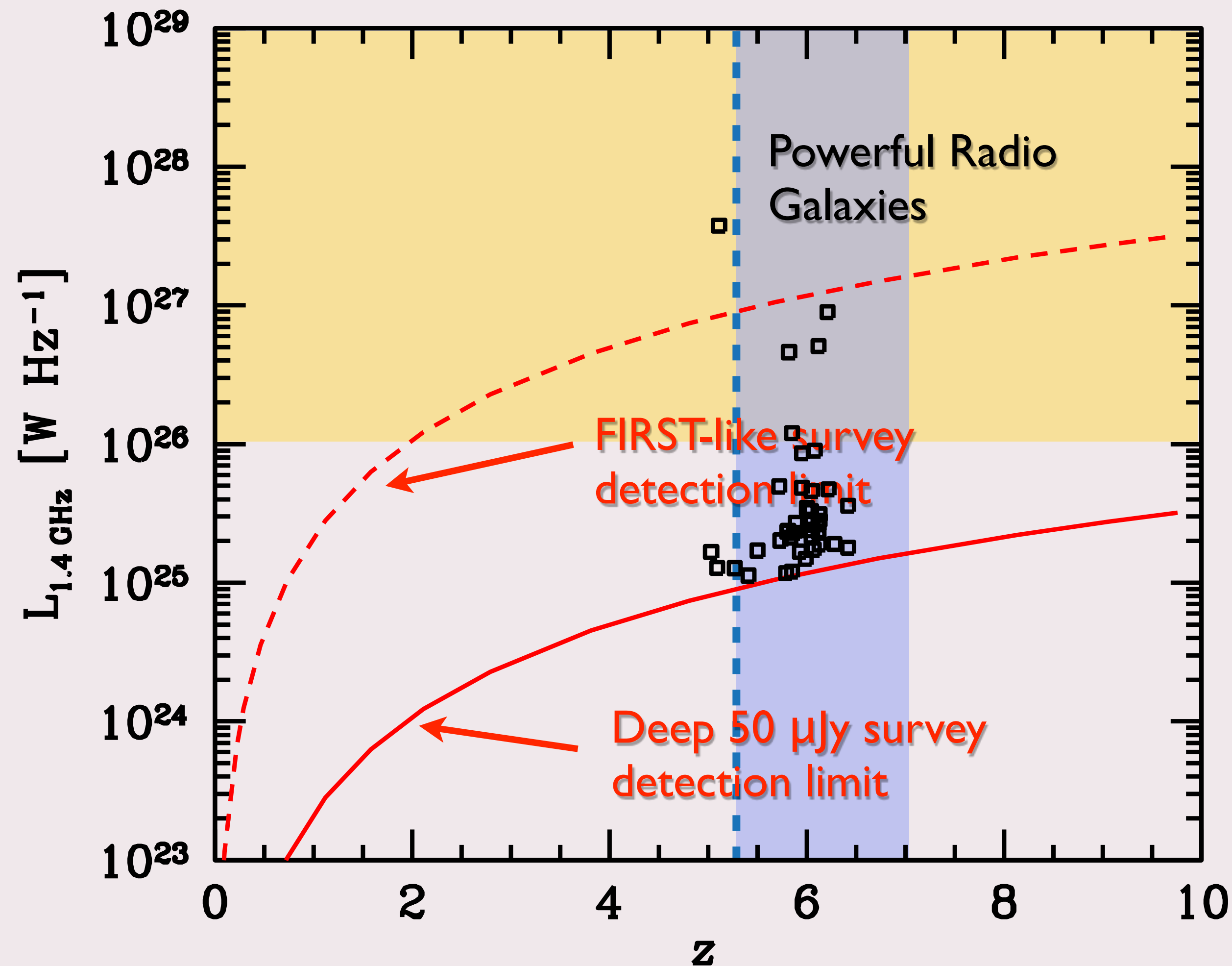
# Do they even exist?



- Observations of high- $z$  AGN:
- Highest  $z$  QSO from UKIDSS ( $z=7.09$ , Mortlock+, II)
  - Over 30 QSOs known at  $6 < z < 7$  (Fan, 2012)



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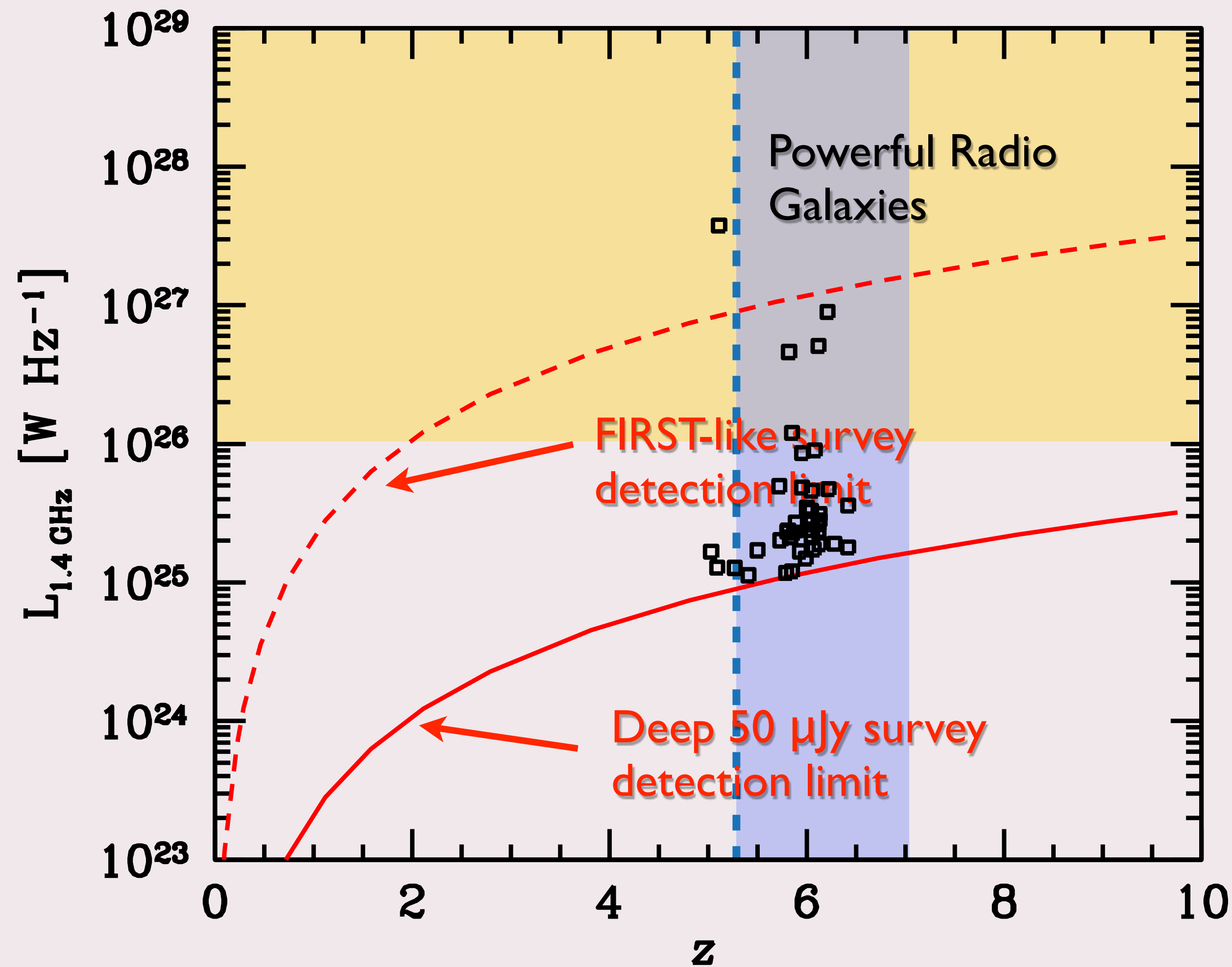
$\Rightarrow 10^8 - 10^9 M_{\odot}$  SMBHs

Powerful AGN *do exist* at  $z \sim 7$ : OK

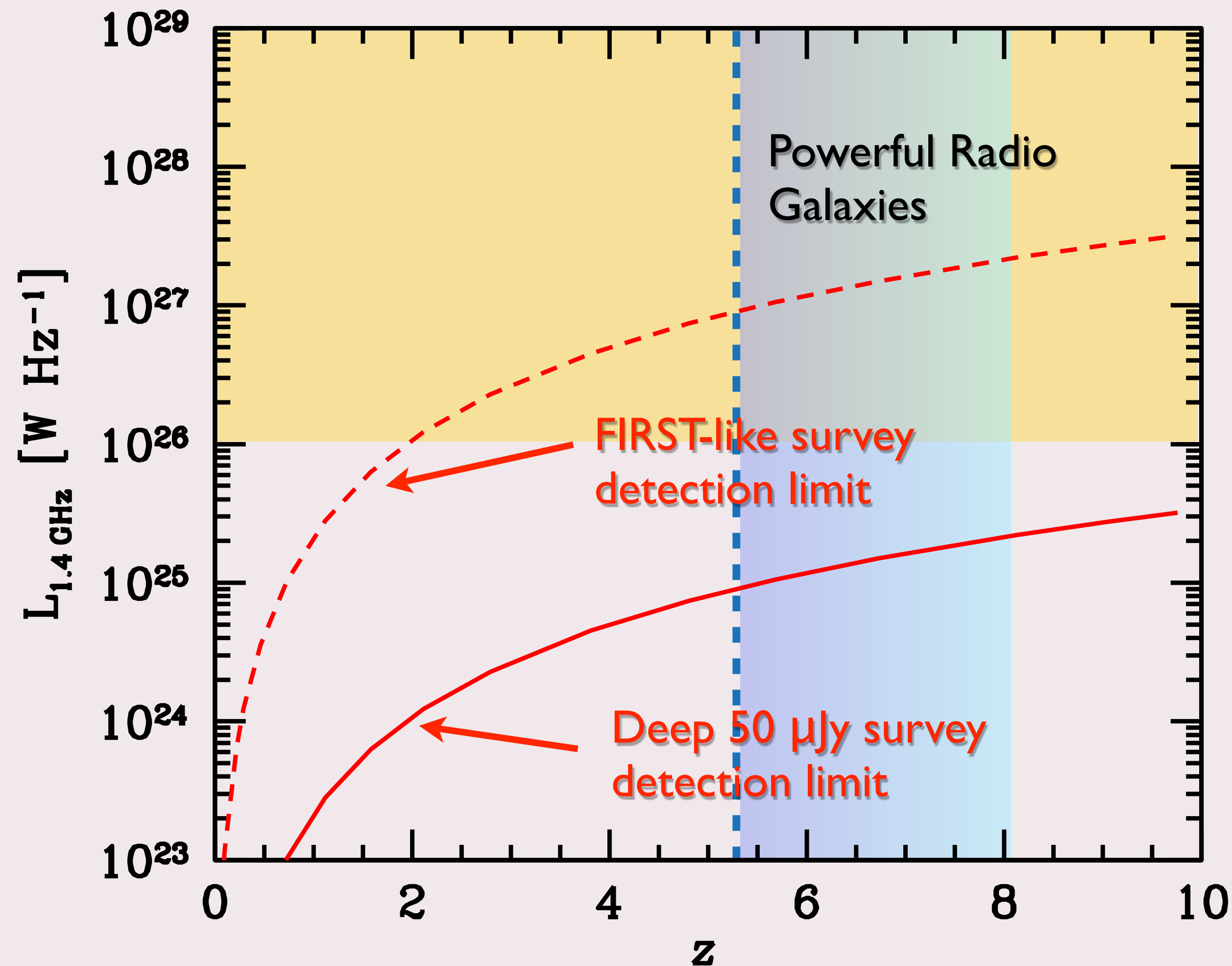


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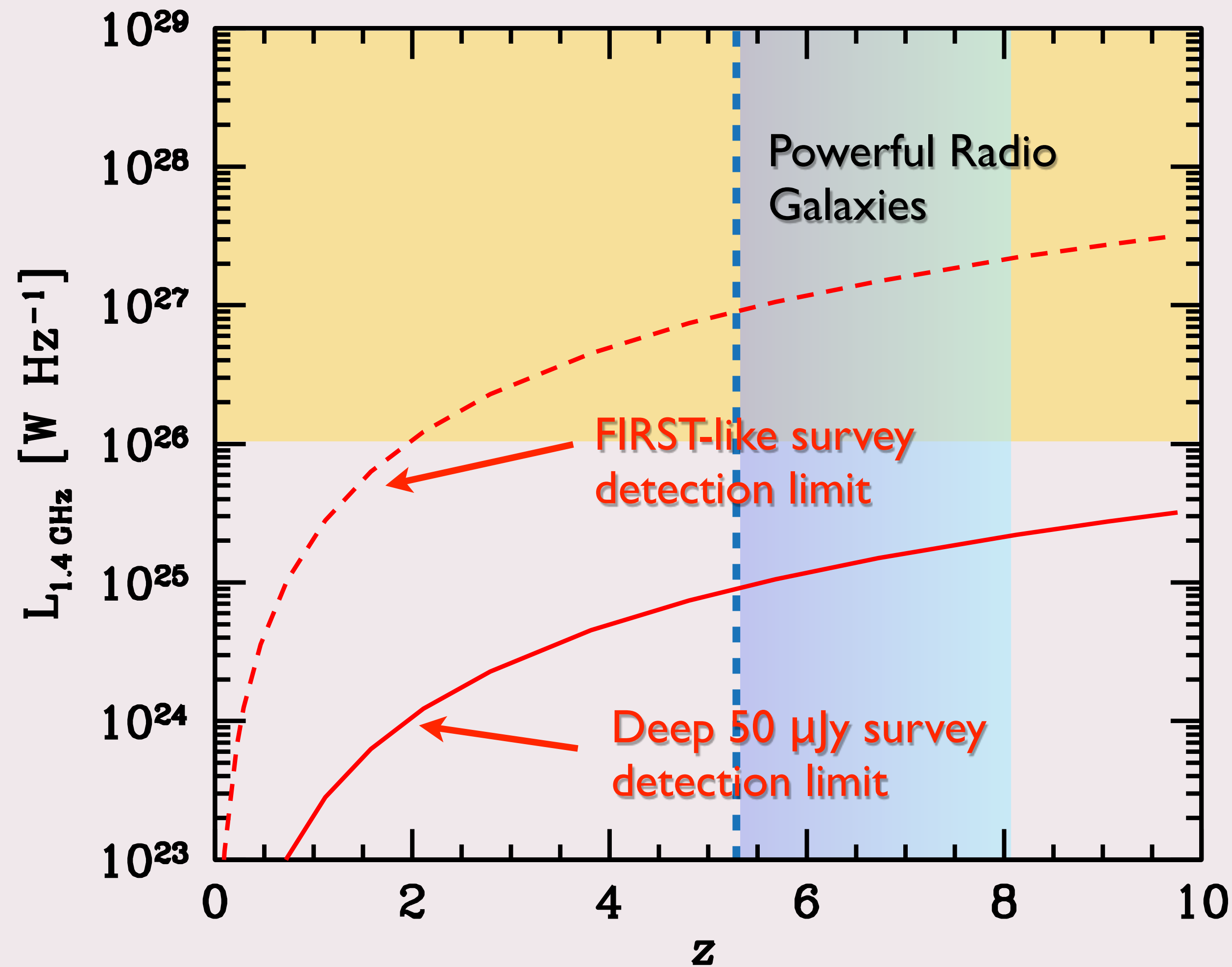
Large pop of SMBHs at  $z \sim 8$   
possible, depending on initial seeds  
(Haiman+, Volonteri+, Johnson+...):

- massive pop III collapse
- DCBH
- runaway collapse in dense stellar clusters at  $z \sim 10-20$

Theory: OK



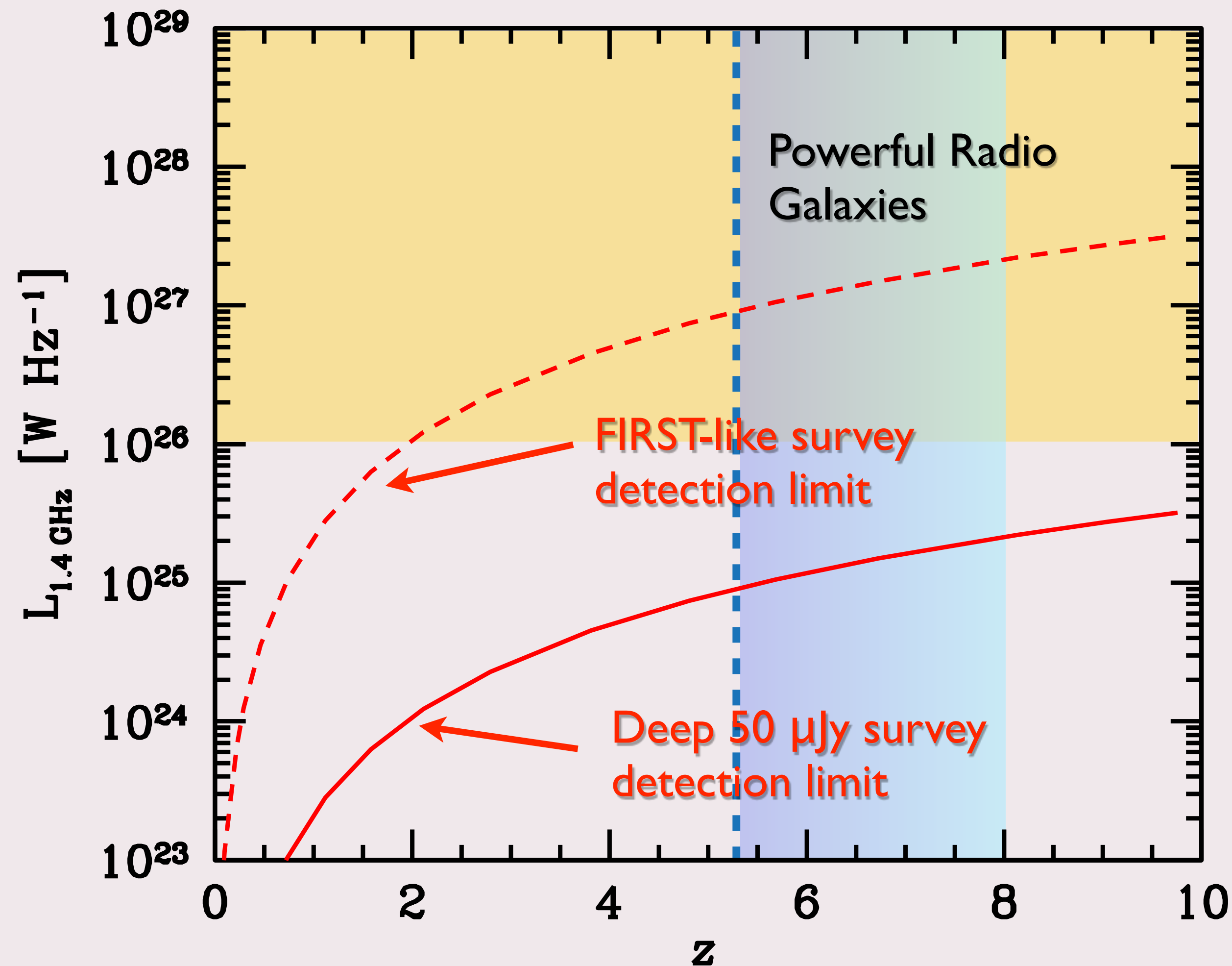
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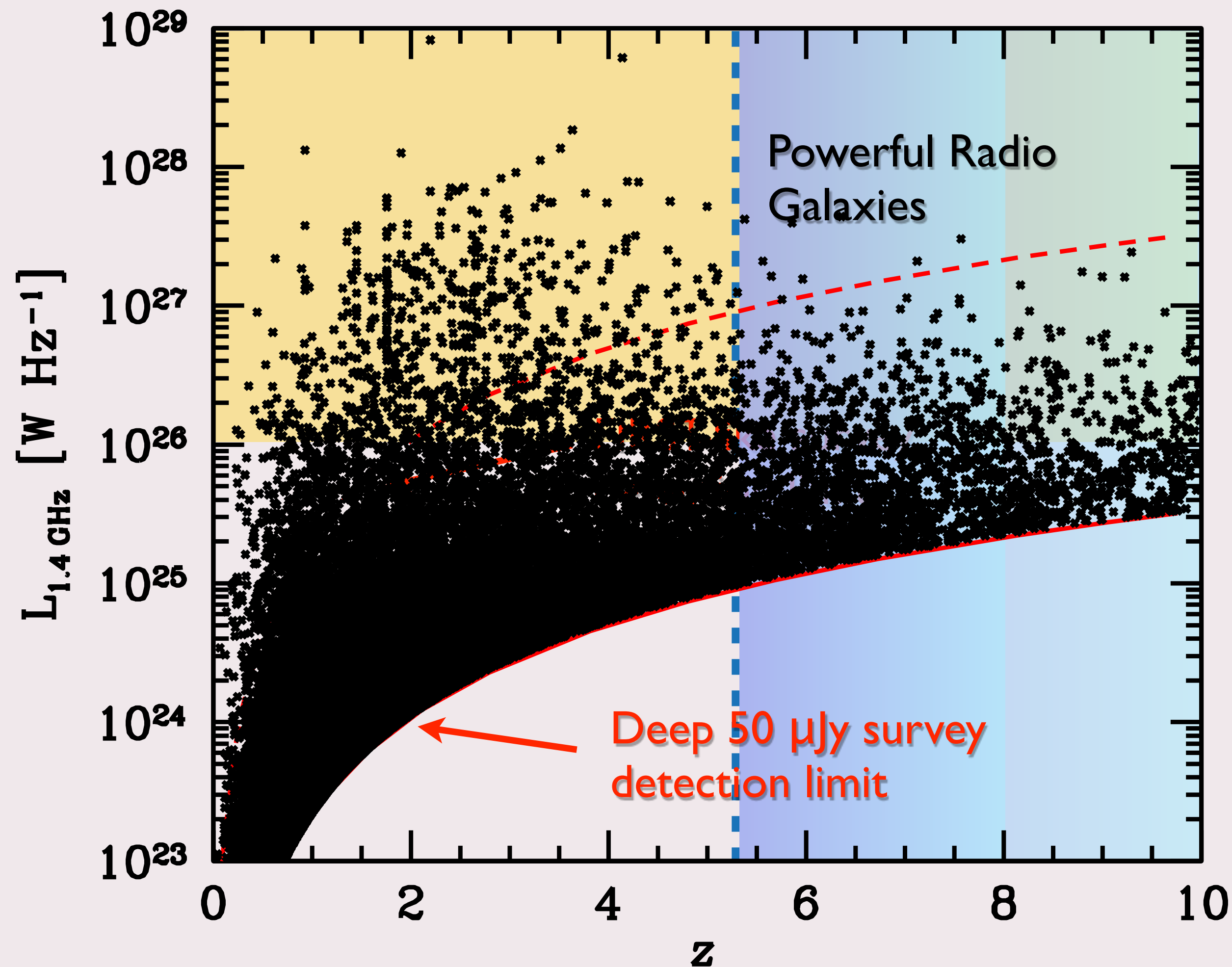
Extrapolate from observations?

SKA Simulated Skies:

- 10 sq deg
- $S_{1.4 \text{ GHz}} > 50 \mu\text{Jy}$



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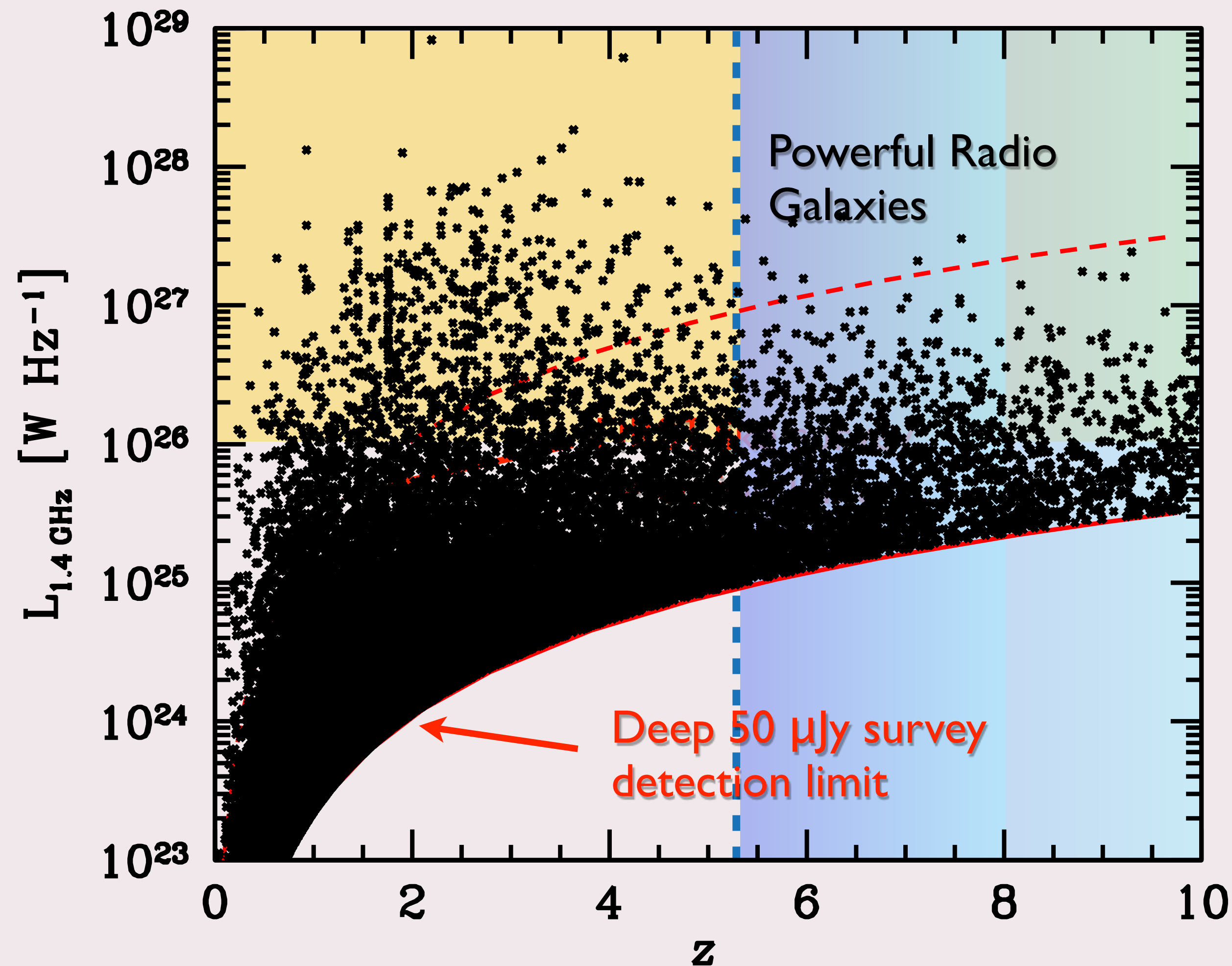
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Radio population evolution: OK

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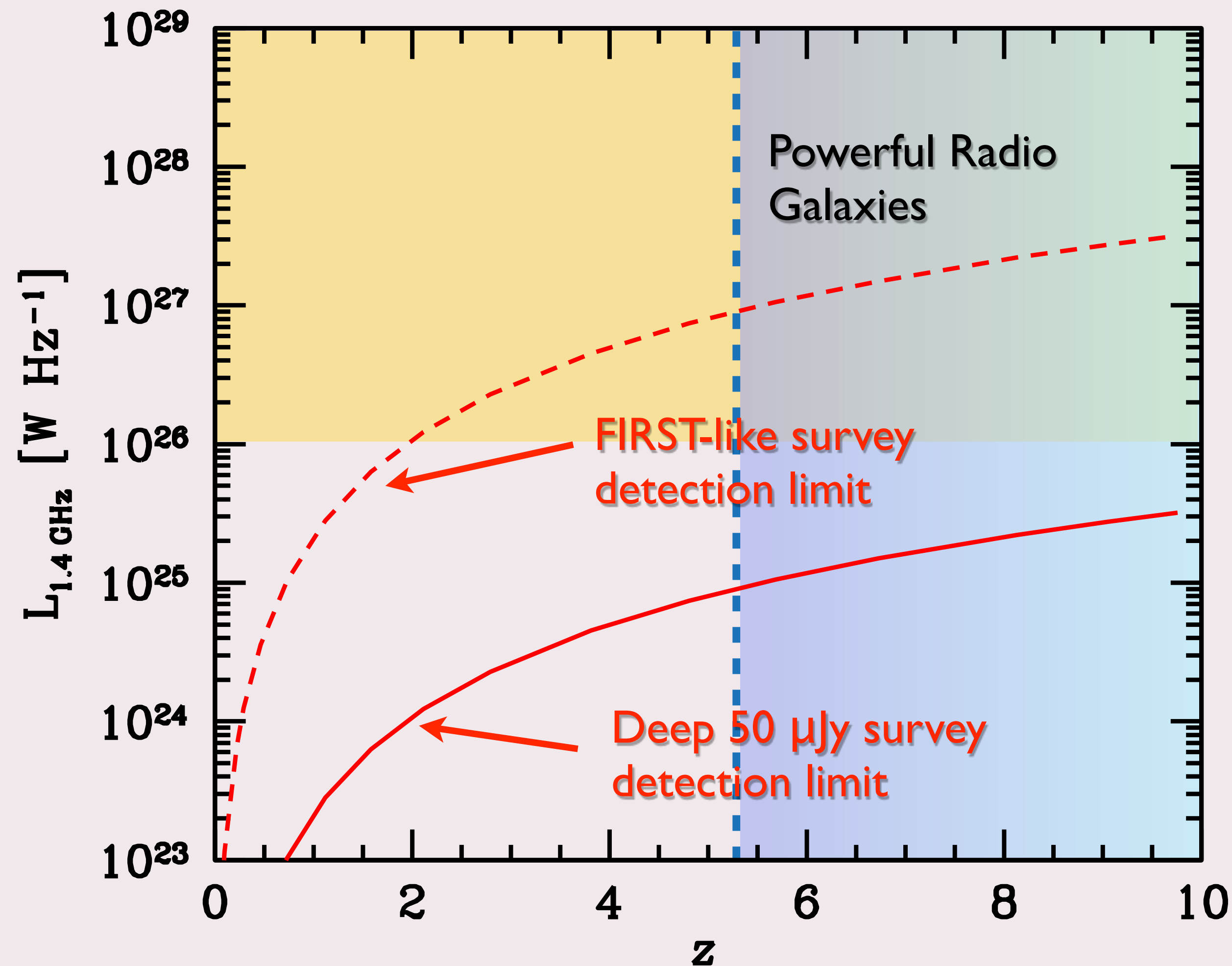
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Theory: OK

Radio population evolution: OK



# Can they be observed?



Powerful AGN *do* exist at  $z \sim 7$ : OK

Theory: OK

Radio population evolution: OK

Increased electron losses to CMB  
- probably not problematic  
(enough energy available; denser environments;...)

CMB losses: probably OK

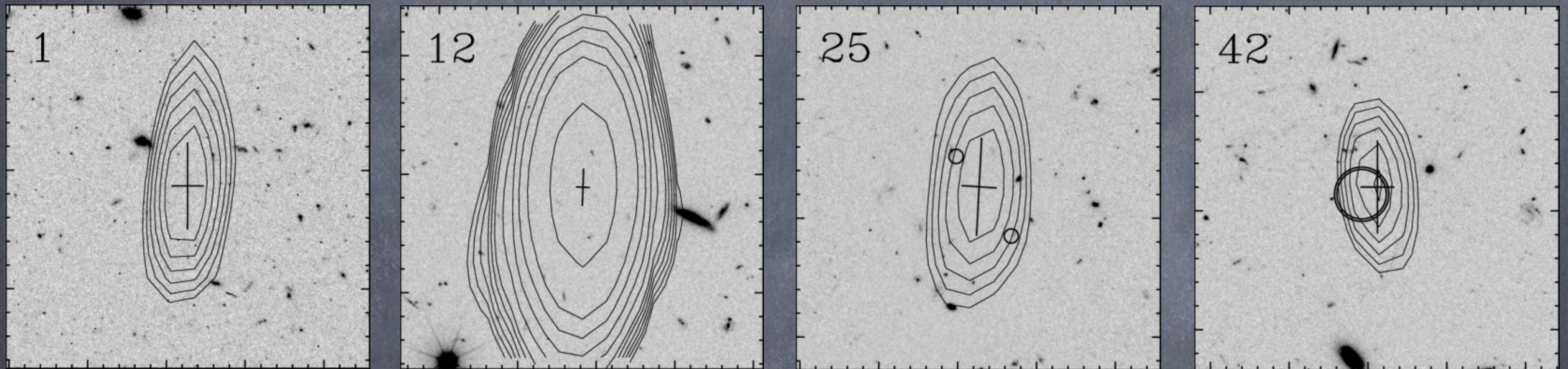
Where are the first radio galaxies?

They (must) have been observed already!



# The currently “unknown” radio panorama

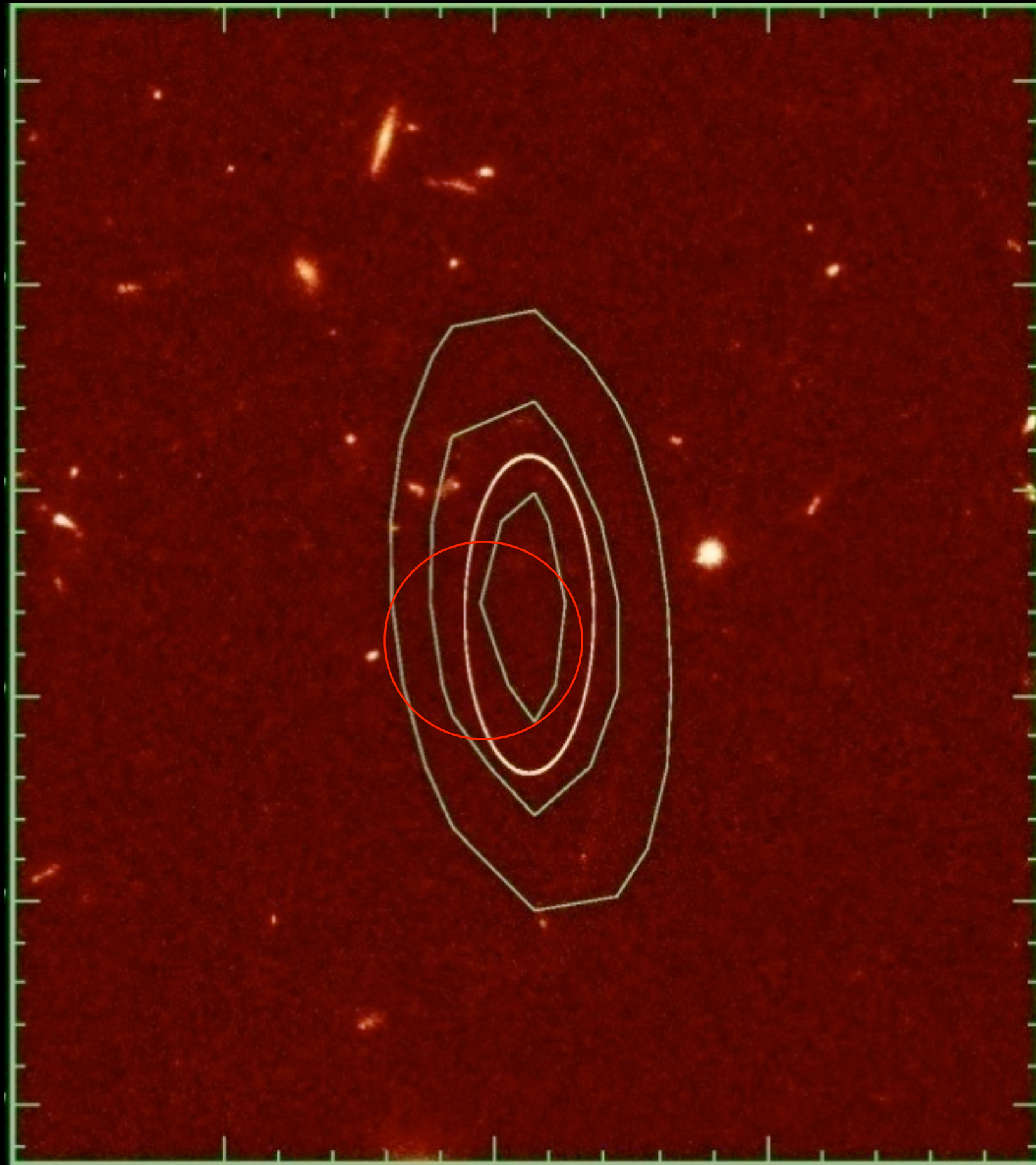
- RGs (including USS) undetected in the optical to very deep levels



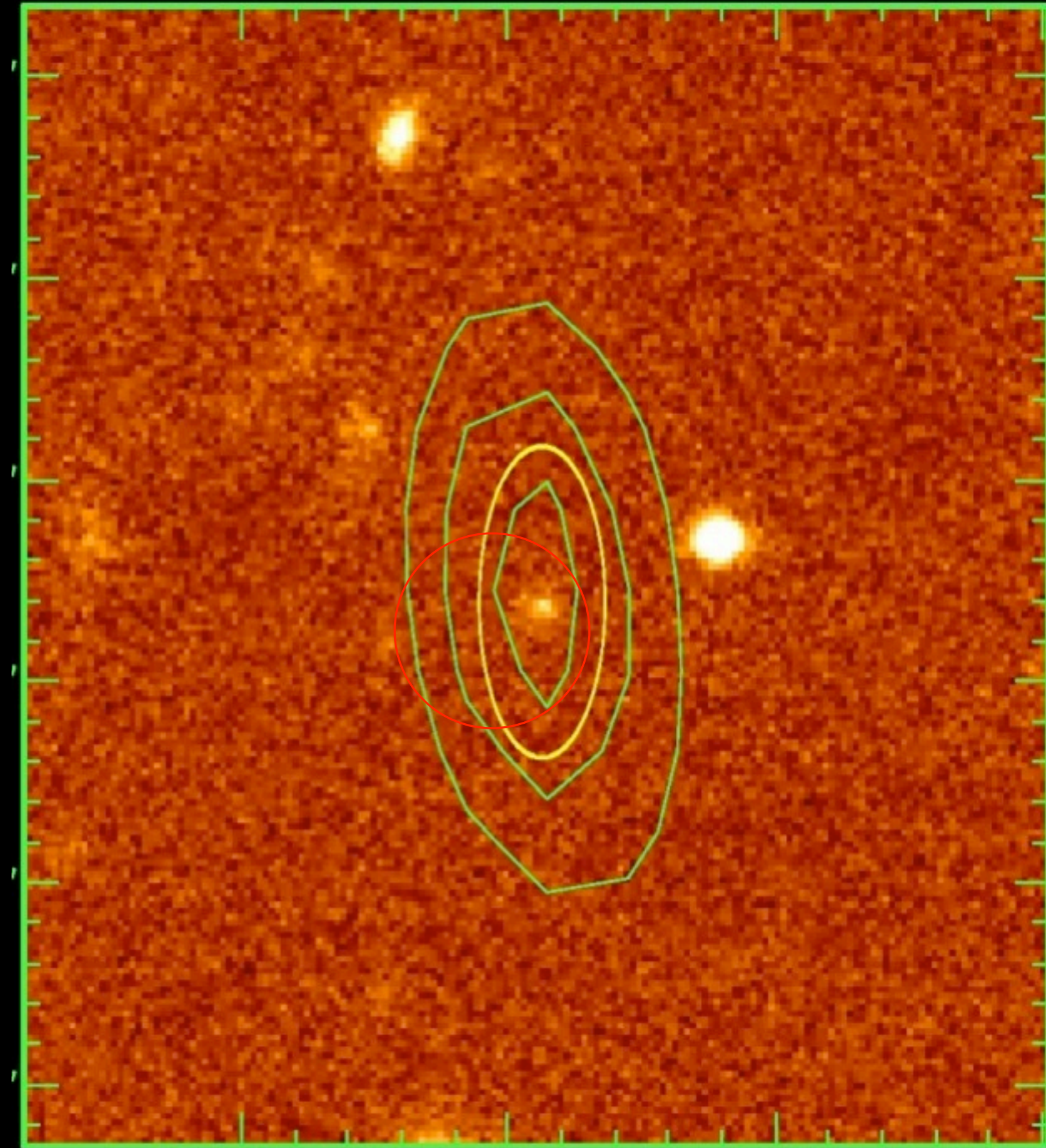
(e.g., CDFS, Lockman Hole: Afonso et al 2006; 2011)



# An extreme ERO AGN



VLT:  $J_{AB}, K_{AB} = 27.2, 23.1$

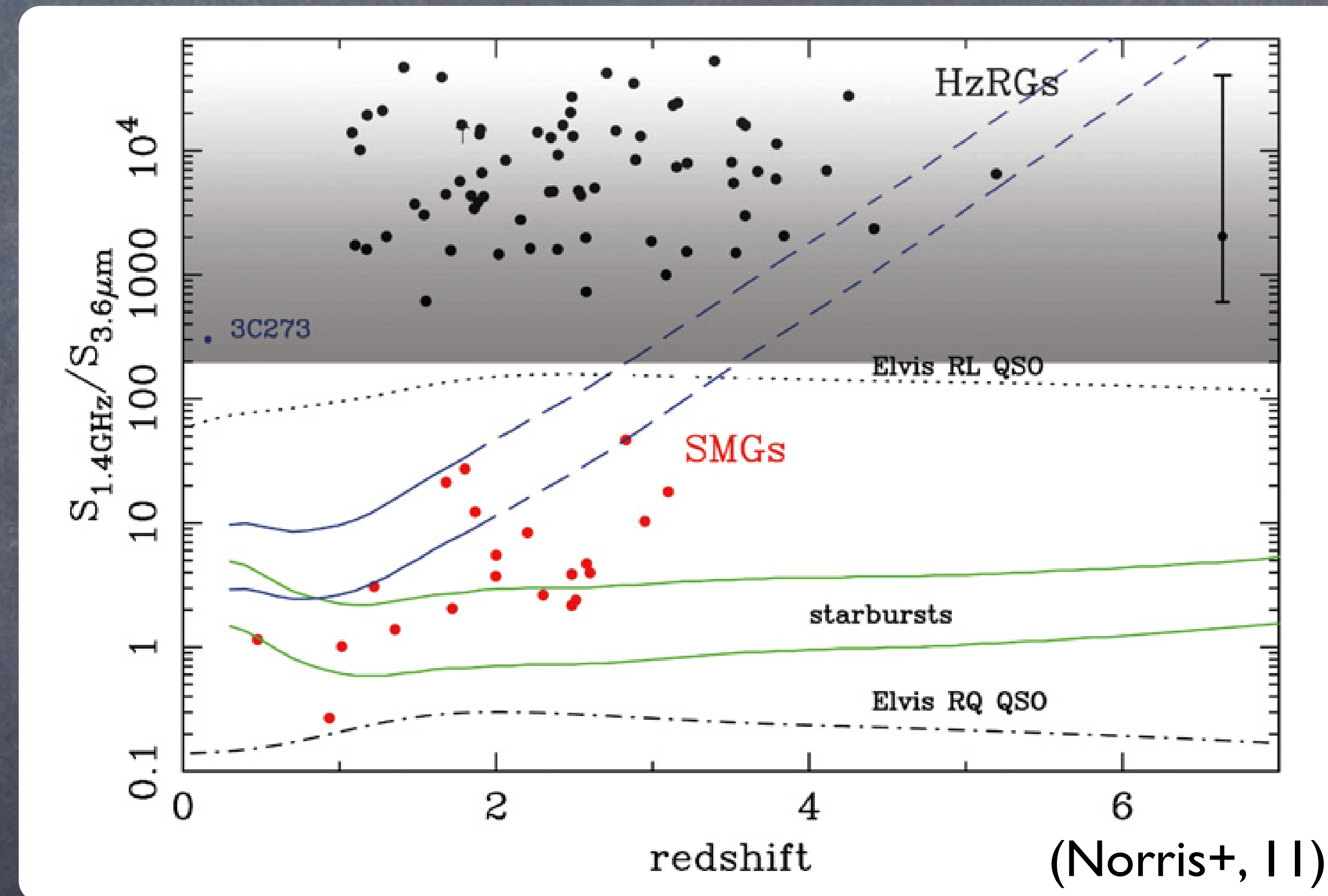


$z-K > 4.5, J-K = 4.1$



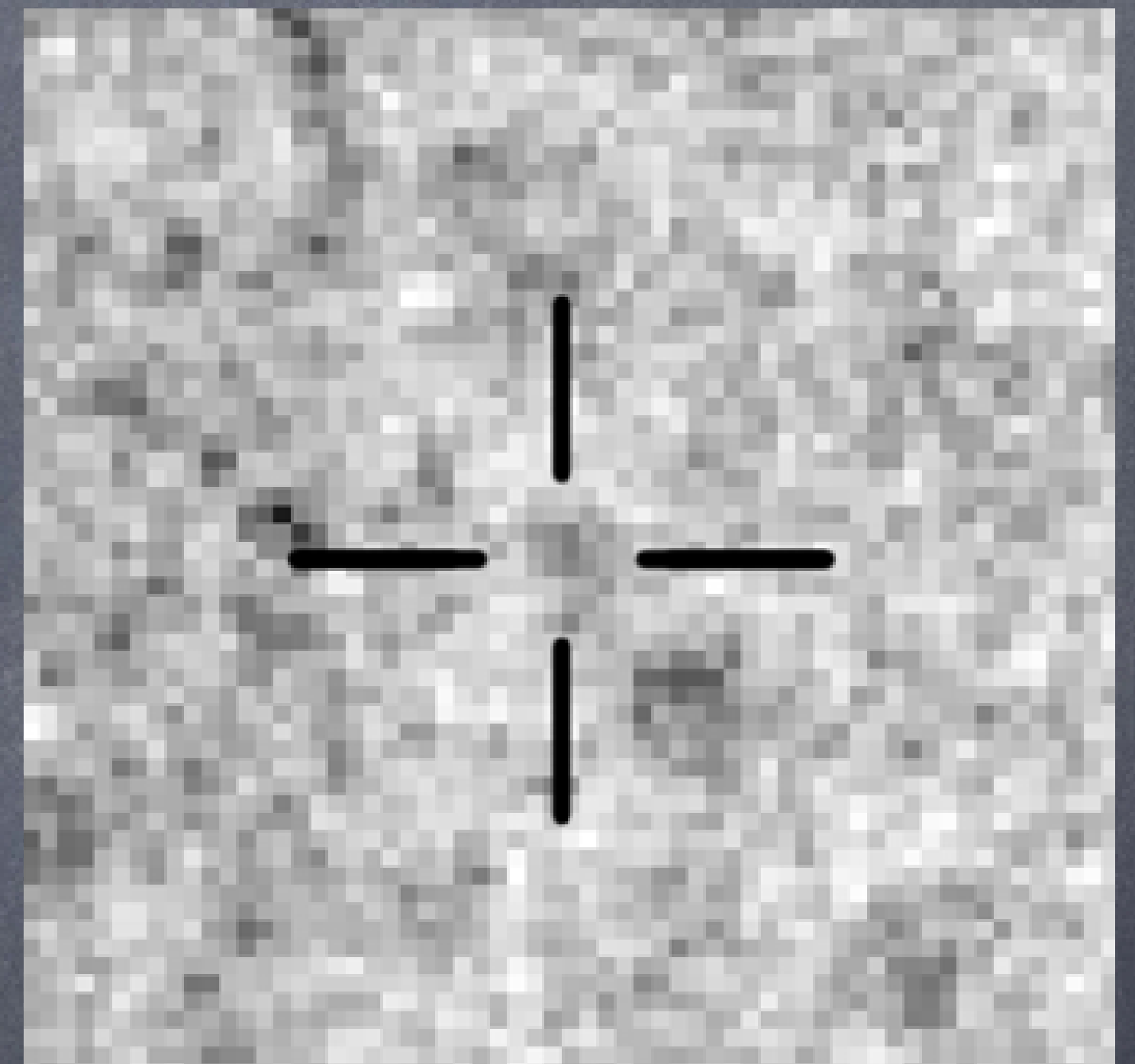
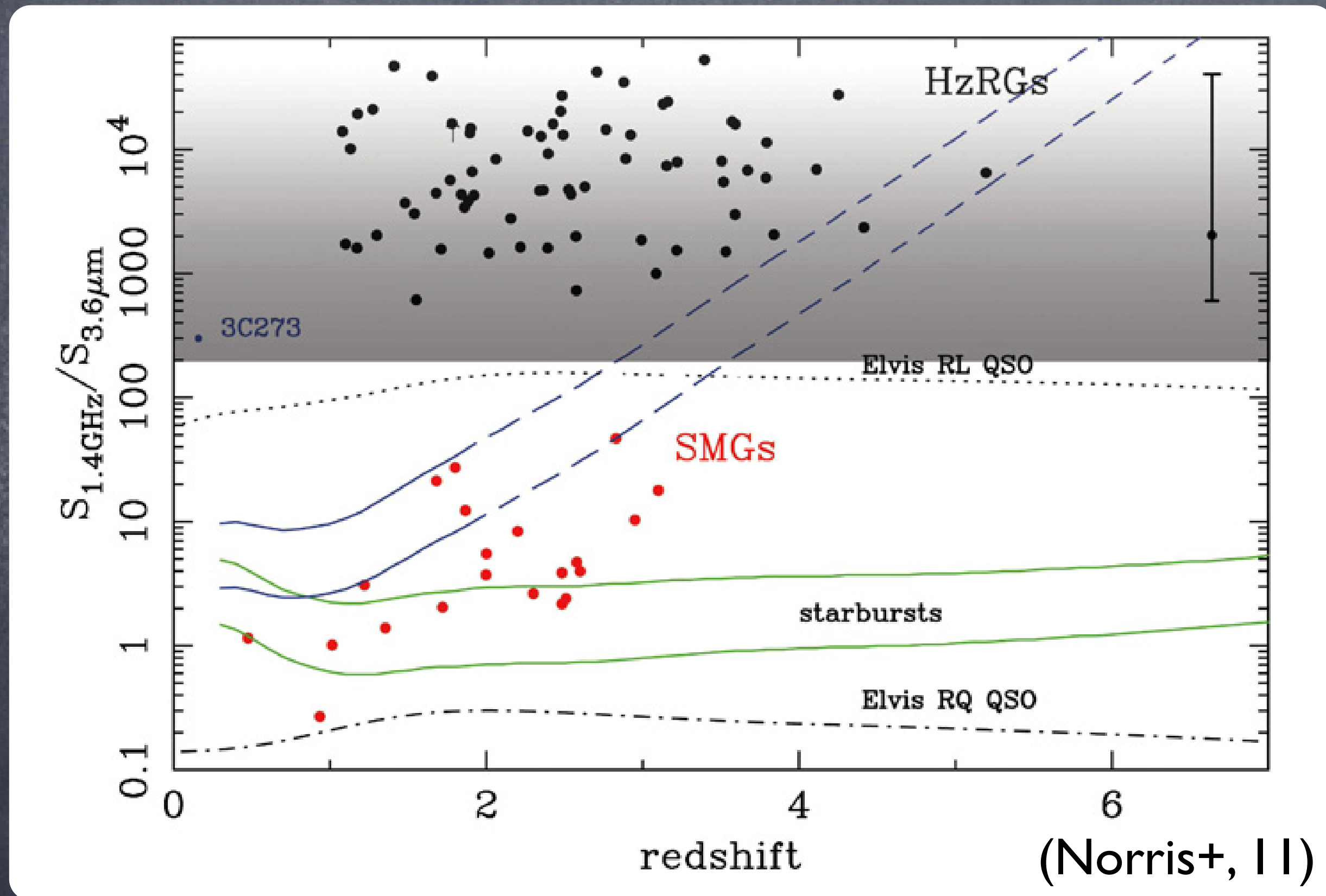
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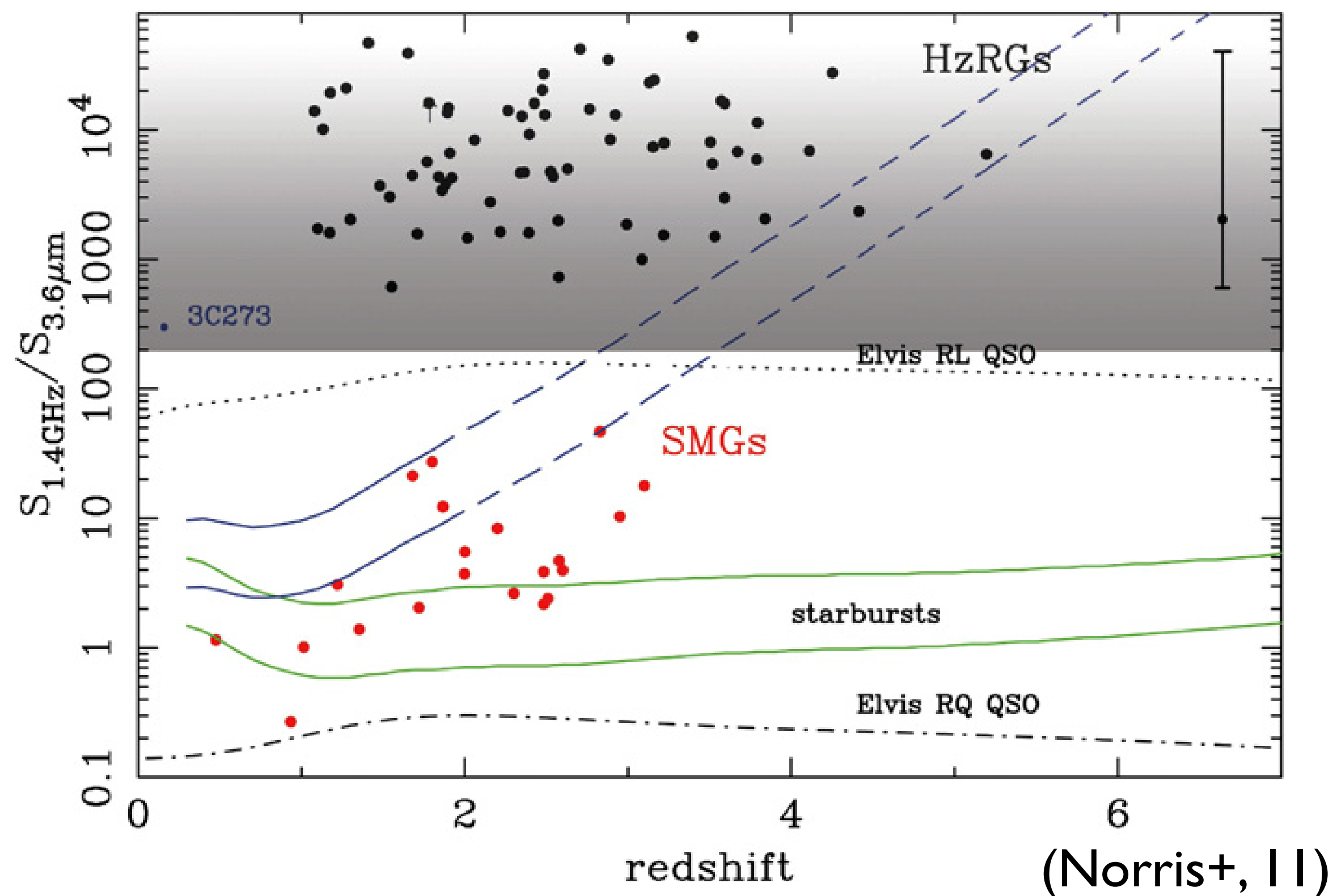


IRAC stack of 39 sources  
 $S_{1.4\text{GHz}} \sim 0.1\text{--}20$  mJy



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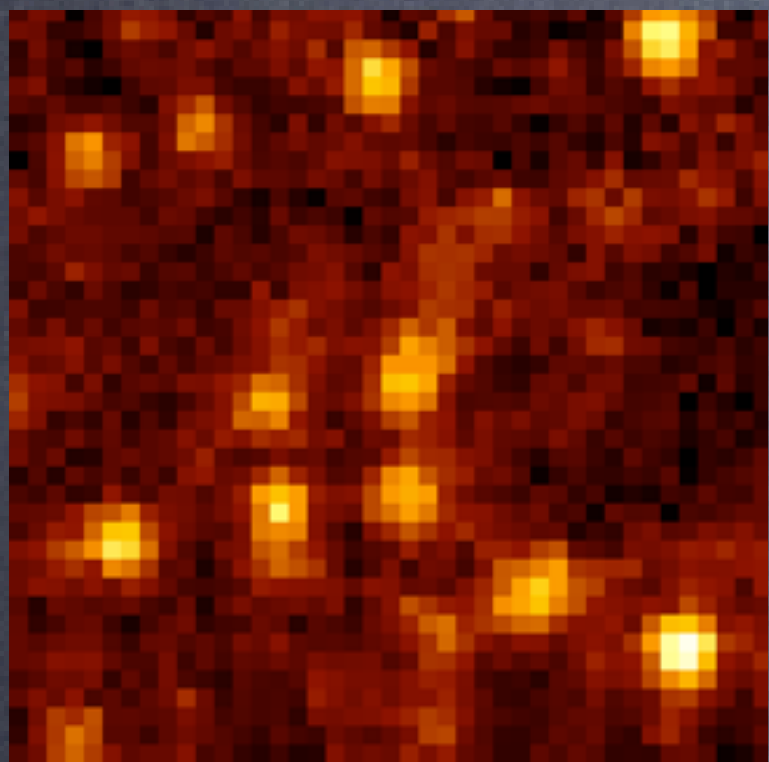


- Possibly, “the most efficient method for selecting high- $z$  radio sources” (Ker+, 2012)

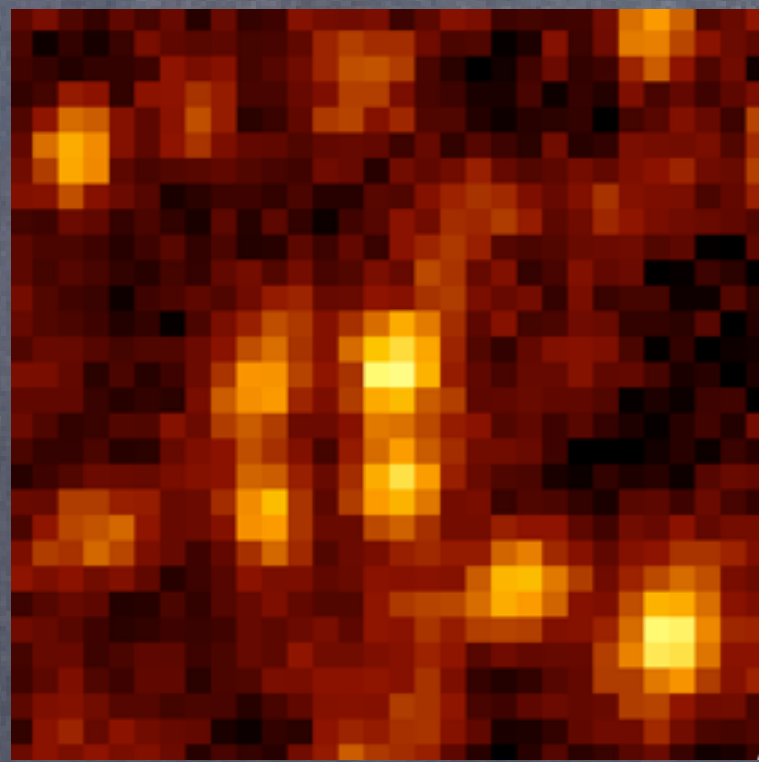
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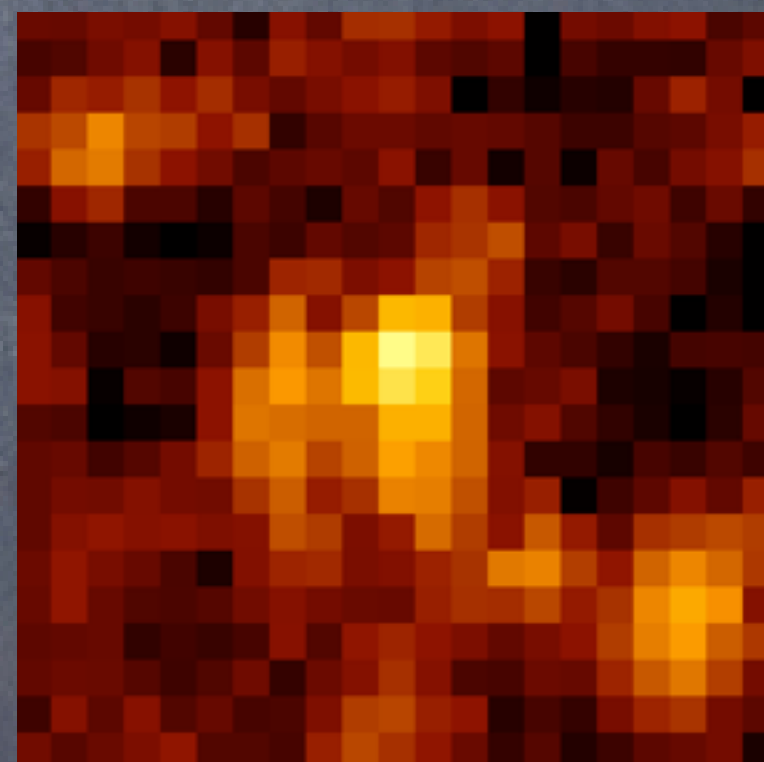
Some are detected with Herschel, with interestingly raising SED slopes in the FIR:



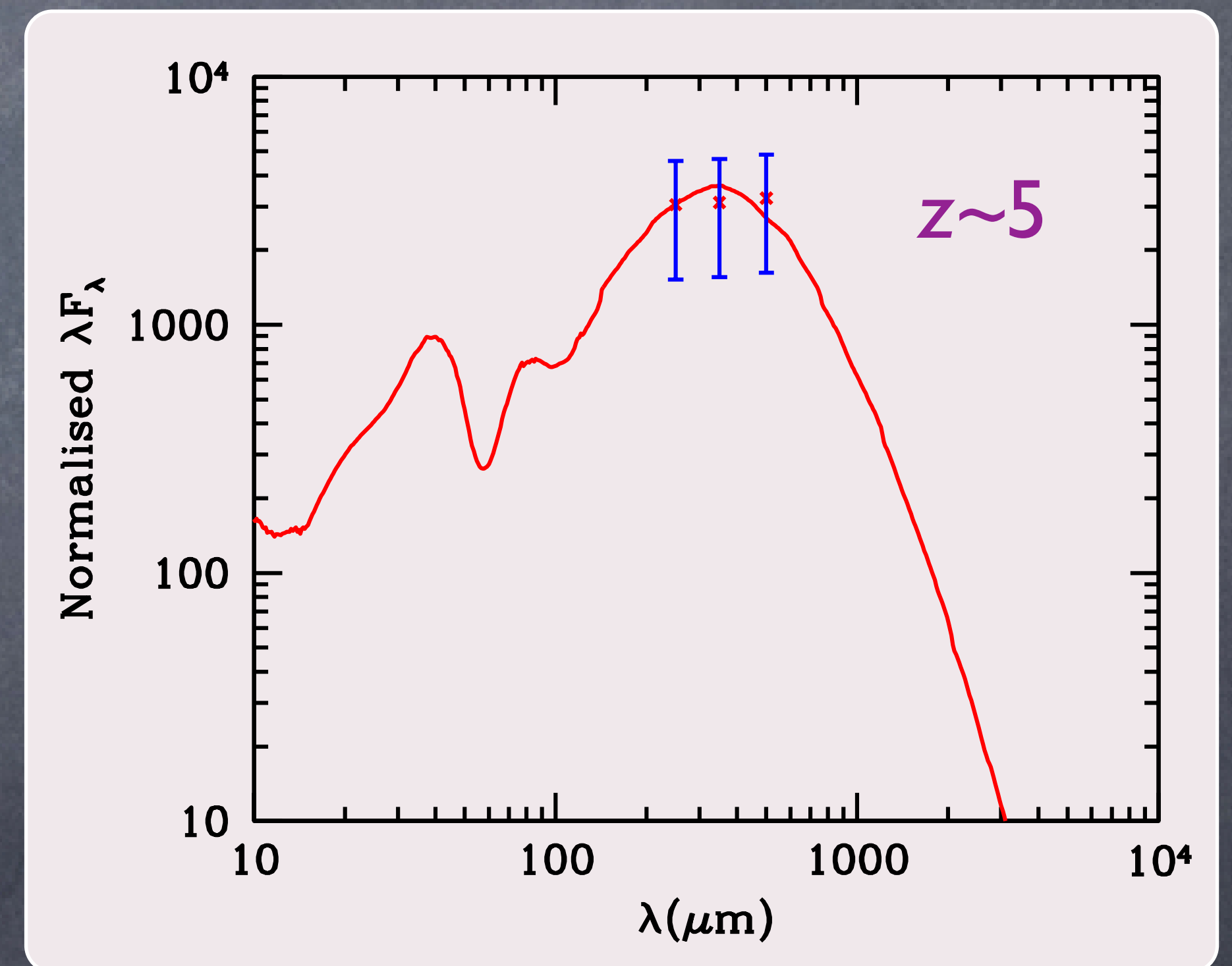
250  $\mu\text{m}$



350  $\mu\text{m}$



500  $\mu\text{m}$

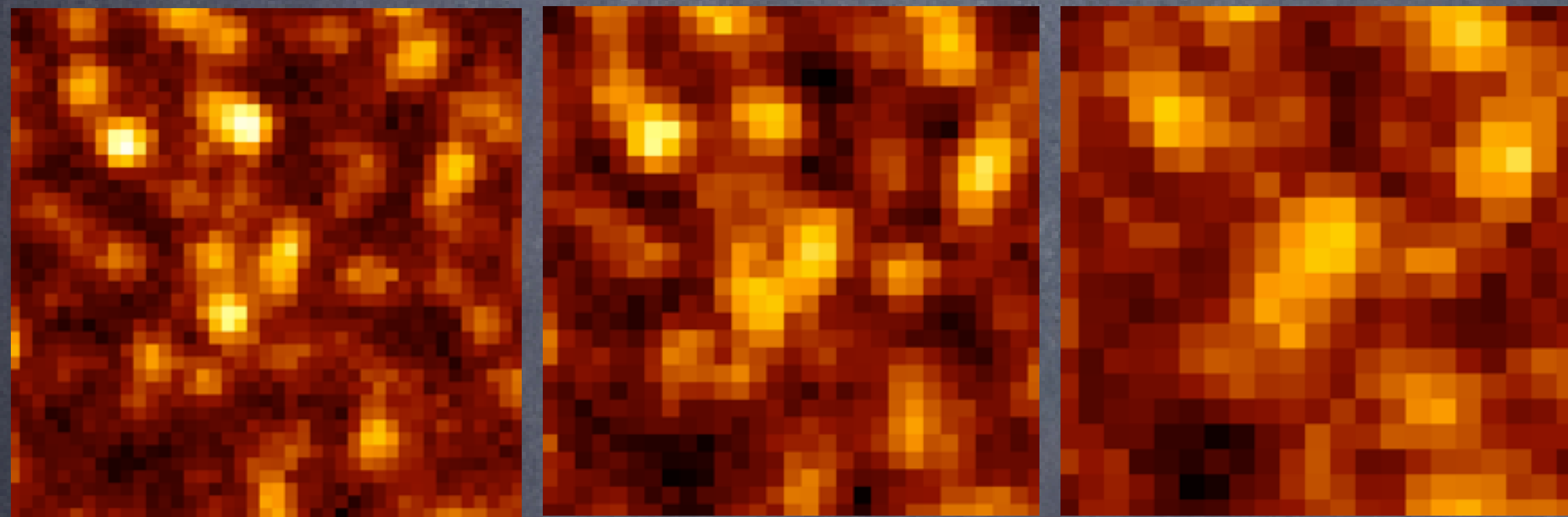




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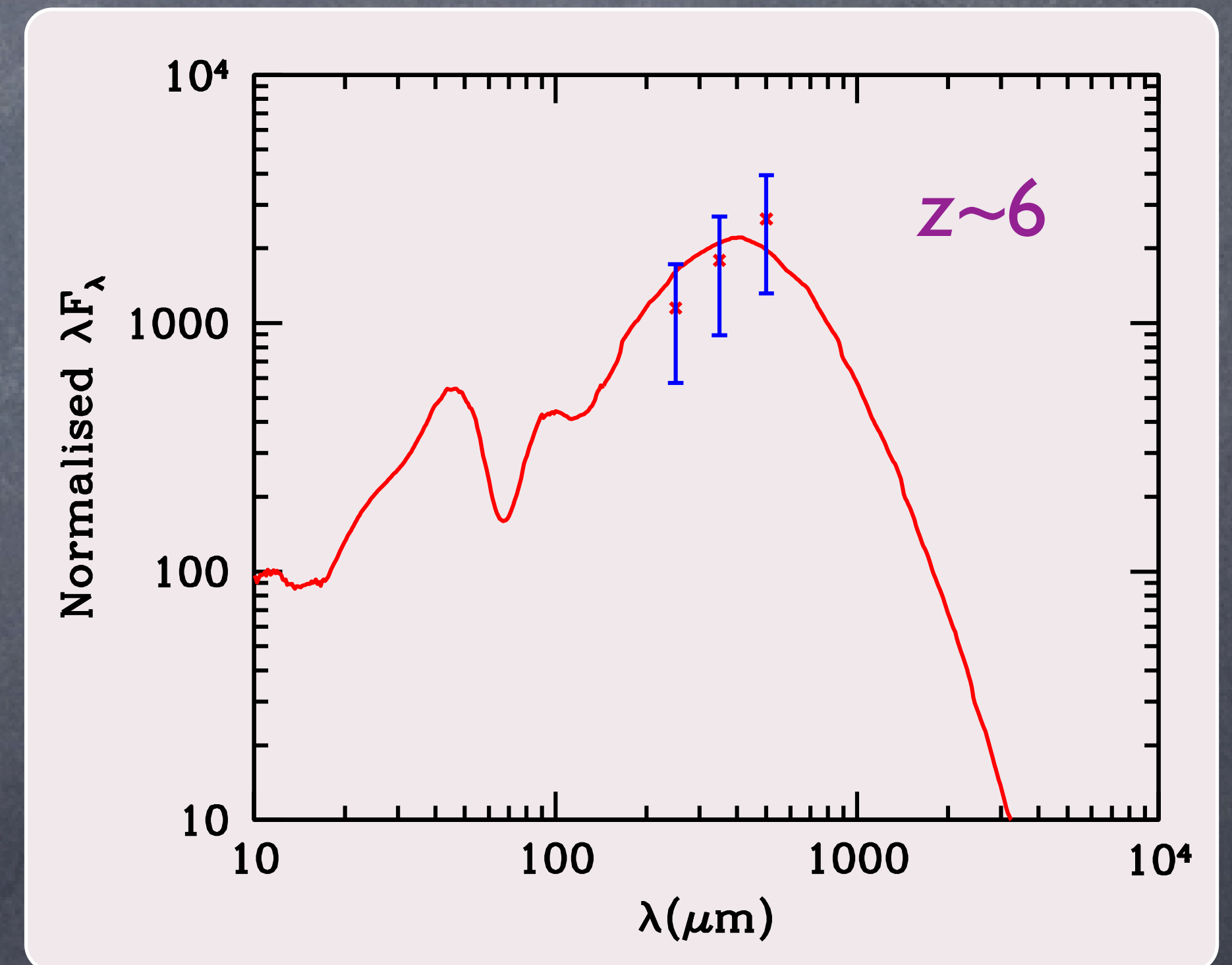
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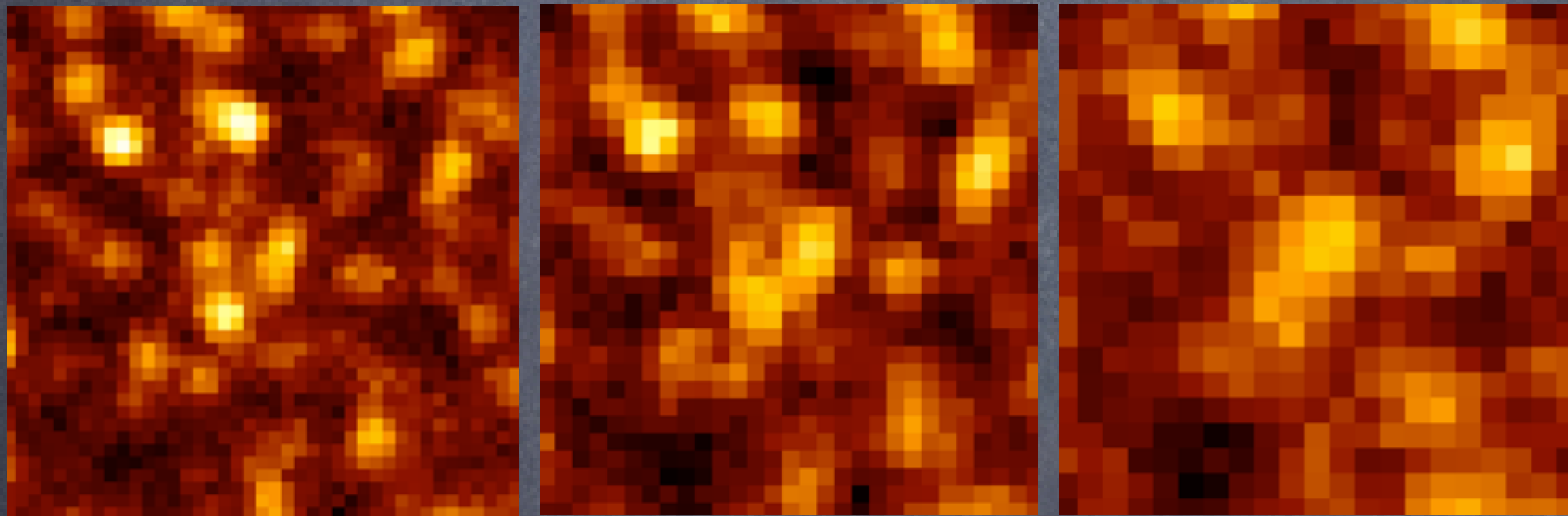
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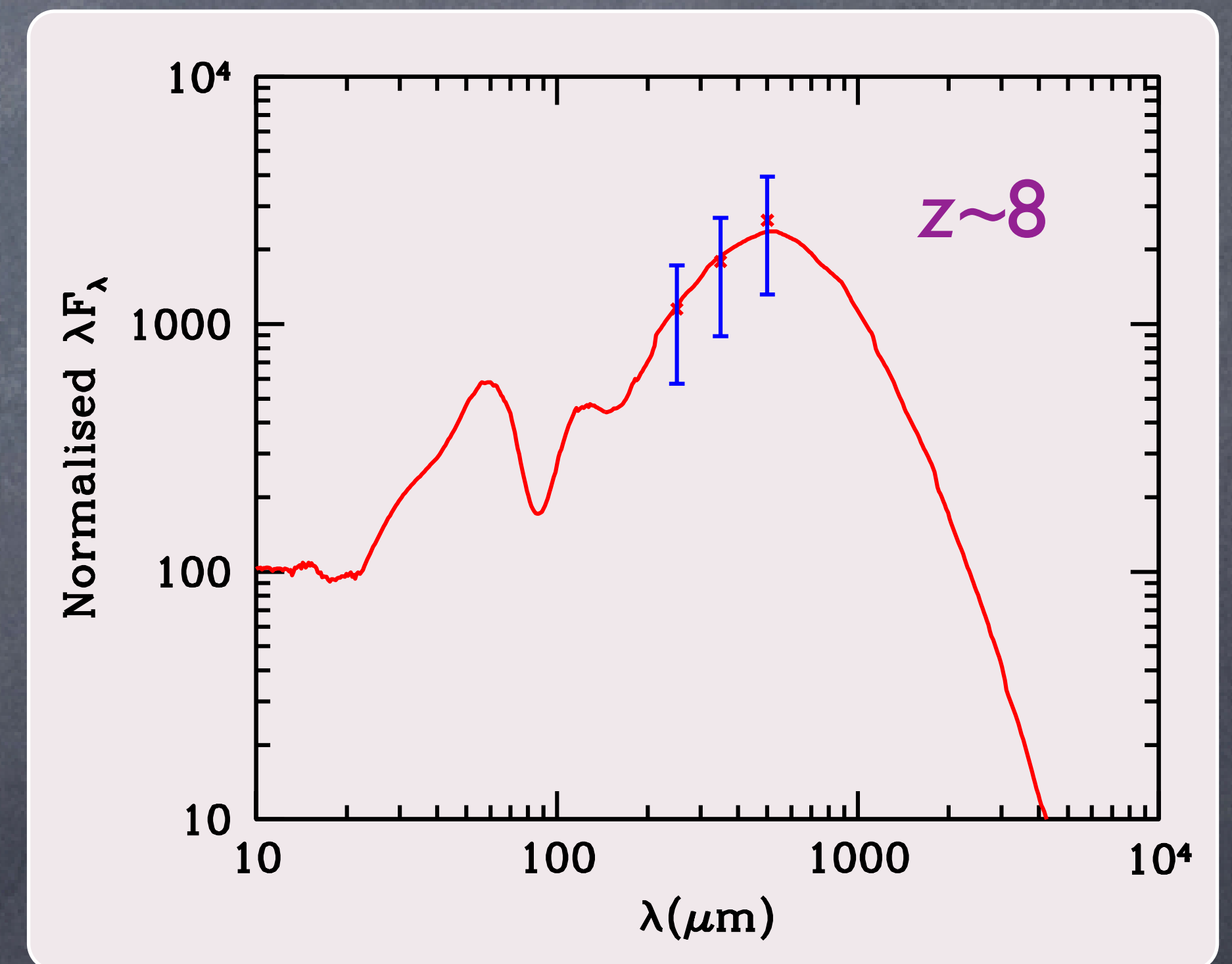
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Ideally (and only) suited for ALMA...



# Upcoming ventures in Deep Radio Probes of the Universe

By 2018, millions of sources from EMU and WODAN (and LOFAR...), over the entire sky!



Evolutionary Map of the Universe  
(ASKAP); Norris+, 11



Westerbork Observations of the Deep  
APERTIF Northern-Sky; Rottgering+, 11

=> FULL SKY @ 1.4GHz,  $\sim 15''$  resolution, 10-20  $\mu\text{Jy}$  rms

Cross-match to IR surveys + {ALMA, JWST, EELT}  
will reveal the (radio)  $z \sim 6-10$  Universe.



# In epitome...

## Where are the first radio galaxies?

- » Radio Surveys have probably already revealed tens of  $z > 6$  AGNs (to  $z \sim ??$ ).
- » Essential to match next generation of radio surveys with deep multiwavelength (IR) follow-up.
- » Confirmation only(?) possible with ALMA, JWST, EELT.



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